

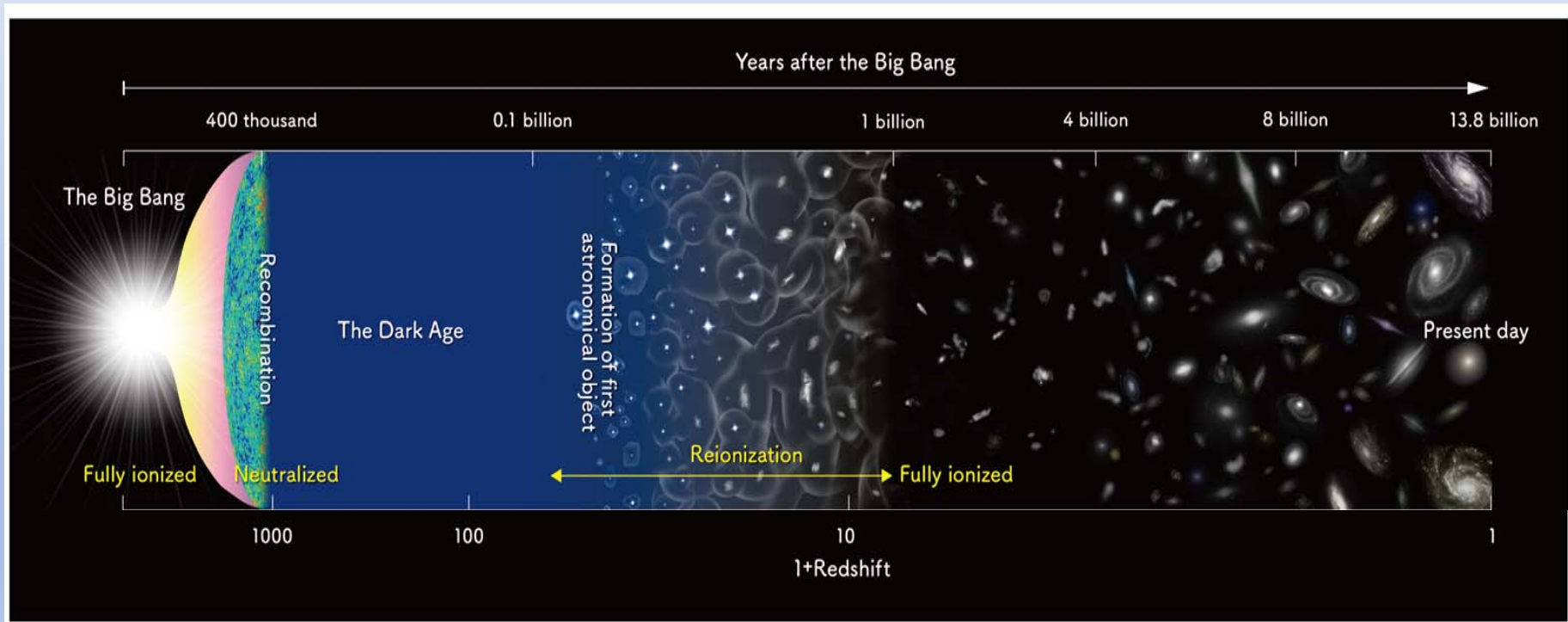
The Spectrum Analysis of a New Quasar at $z=6.61$

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Supervisors : Tomotsugu Goto, Youichi Ohyama

Why High Redshift ($z > 6$) Quasars Are Important?

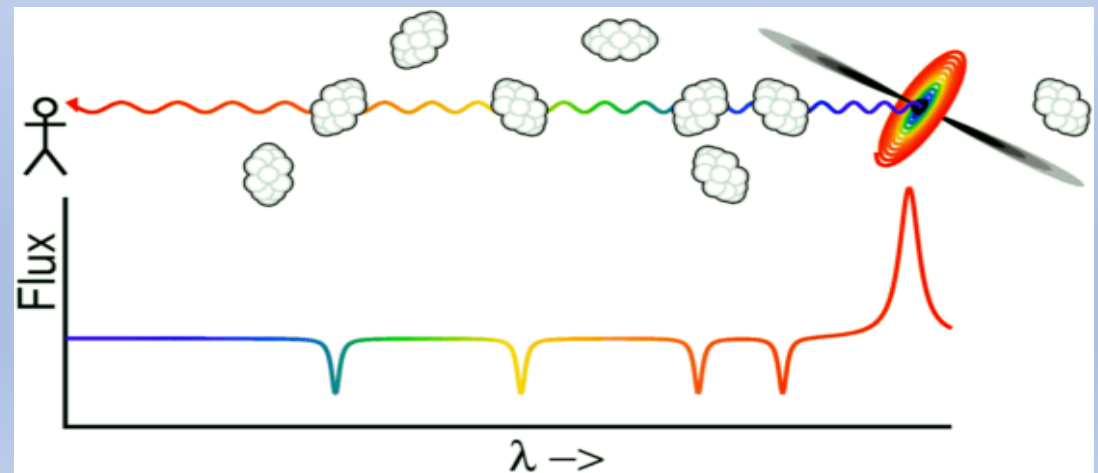
- Reionization epoch



(Credit: NAOJ)

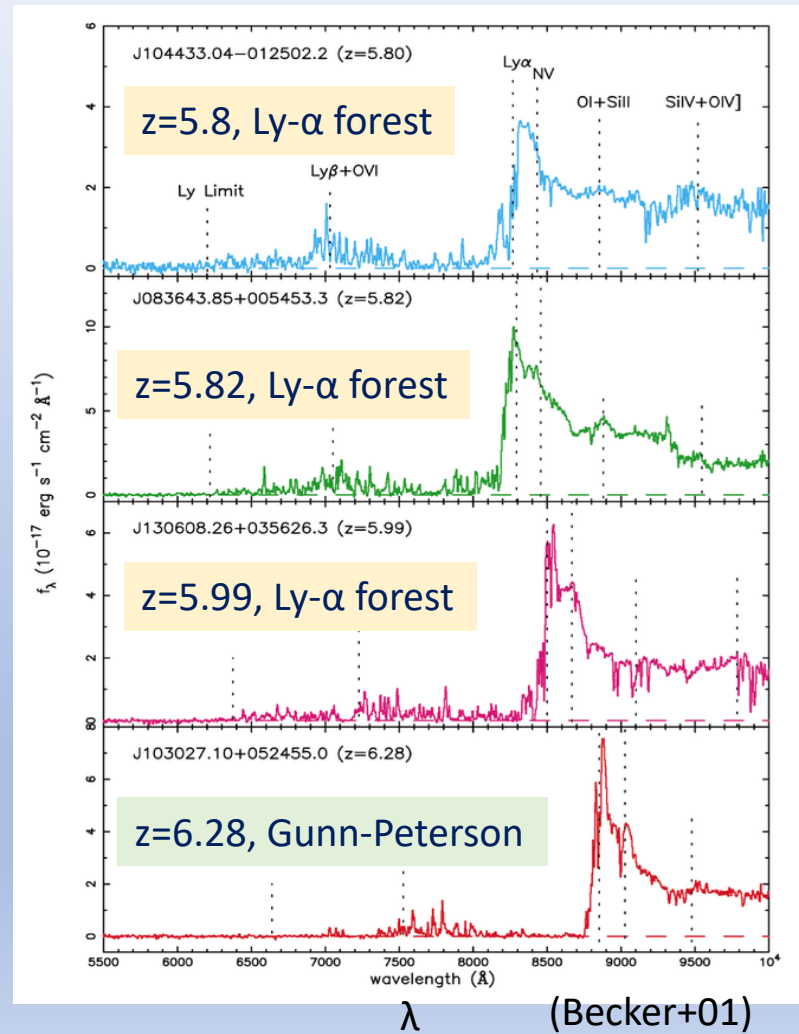
What Can the Spectrum of the High Redshift Quasar Tell Us About the Reionization Epoch?

- The information along the line of sight
- Larger scale-neutral hydrogen
 - Lyman-alpha forest or Gunn-Peterson effect
 - Neutral hydrogen fraction

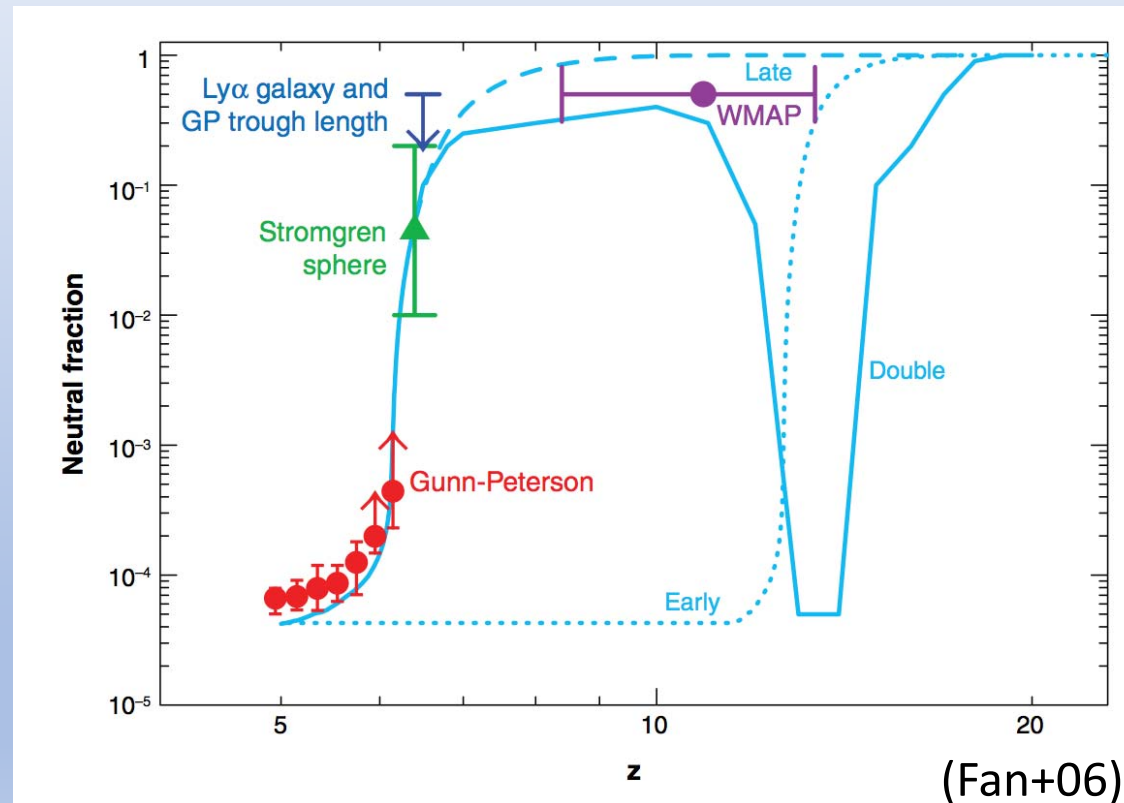


(Credit from <http://www.astro.ucla.edu/~wright/Lyman-alpha-forest.html>)

Reionization Epoch - Neutral Hydrogen Fraction

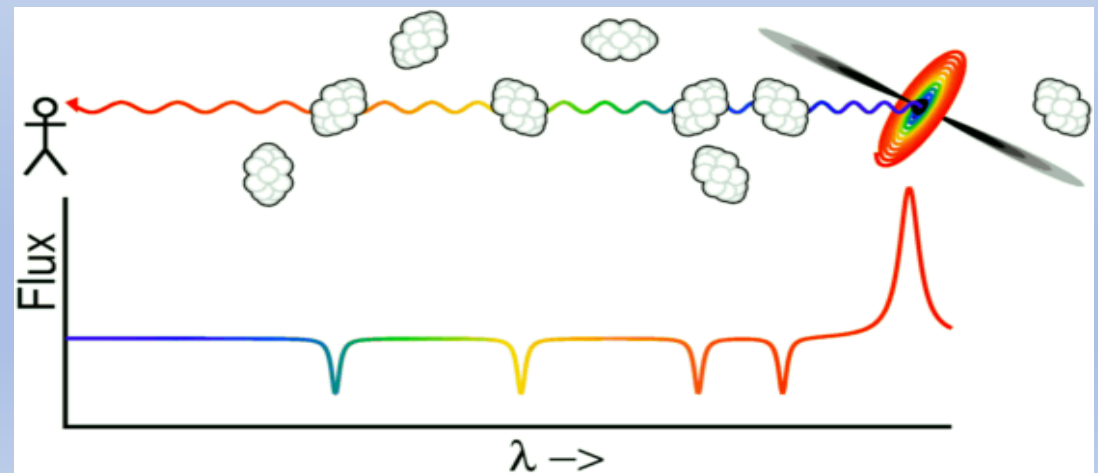


Reionization Epoch - Neutral Hydrogen Fraction



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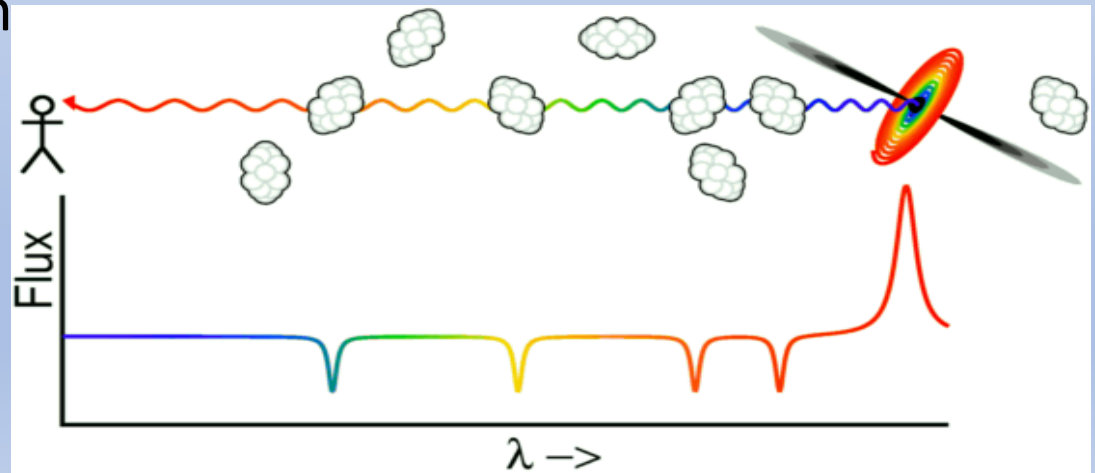
- The information along the line of sight
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What Can the Spectrum of the High Redshift Quasar Tell Us About the Reionization Epoch?

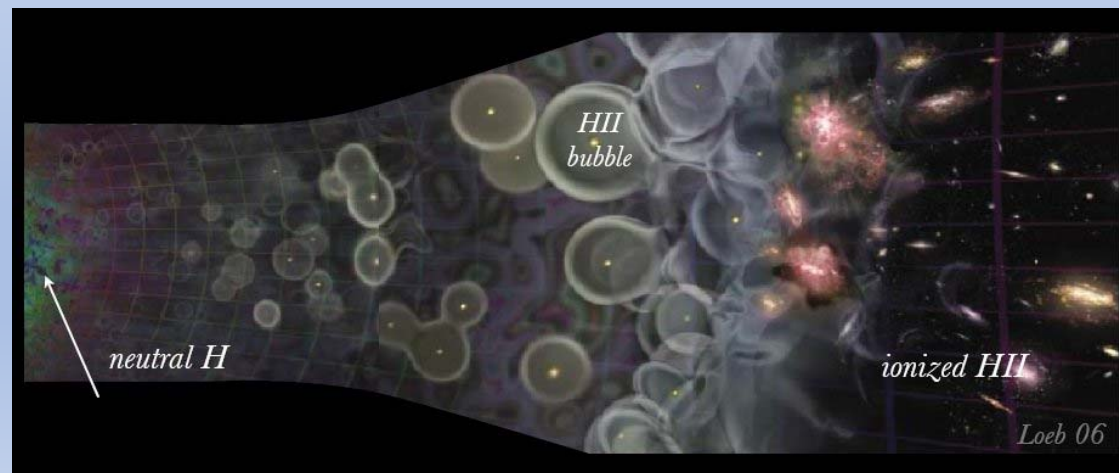
- The information along the line of sight
- Larger scale-neutral hydrogen
 - Lyman-alpha forest or Gunn-Peterson effect
 - Neutral hydrogen fraction
- Smaller scale-ionized hydrogen
 - Near-zone region



(Credit from <http://www.astro.ucla.edu/~wright/Lyman-alpha-forest.html>)

Reionization Epoch - Near Zone Region

- Show how much a QSO can ionize
- Ly- α emission line emits from the quasar
- The “HII bubble” ionized by the quasar
- Transmitted through the bubble
- A few Mpc

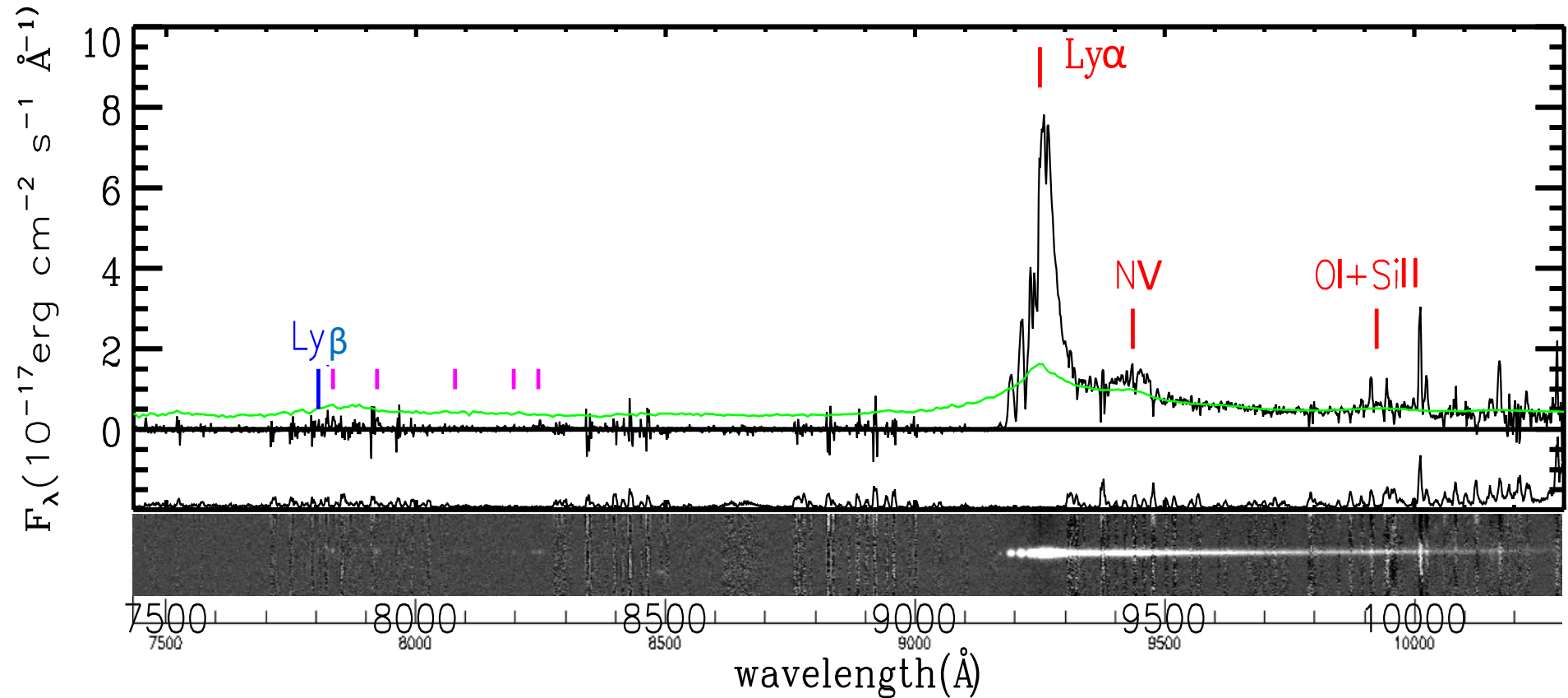


(Loeb+06)

The Quasar Discovered at $z=6.61$ from Pan-STARRS1

- Quasar J006+39
- Subaru-FOCAS spectroscopic observation
 - 2015.11.02, seeing $\sim 0.75''$
 - Slit width: $0.8''$
 - Resolution ~ 600
 - 1000sec exposure X 5 times
 - VPH900 grating, O58 order cut
 - $7434\text{\AA} \sim 10582\text{\AA}$ with 1.48\AA pixel resolution

Quasar J006+39 Spectrum



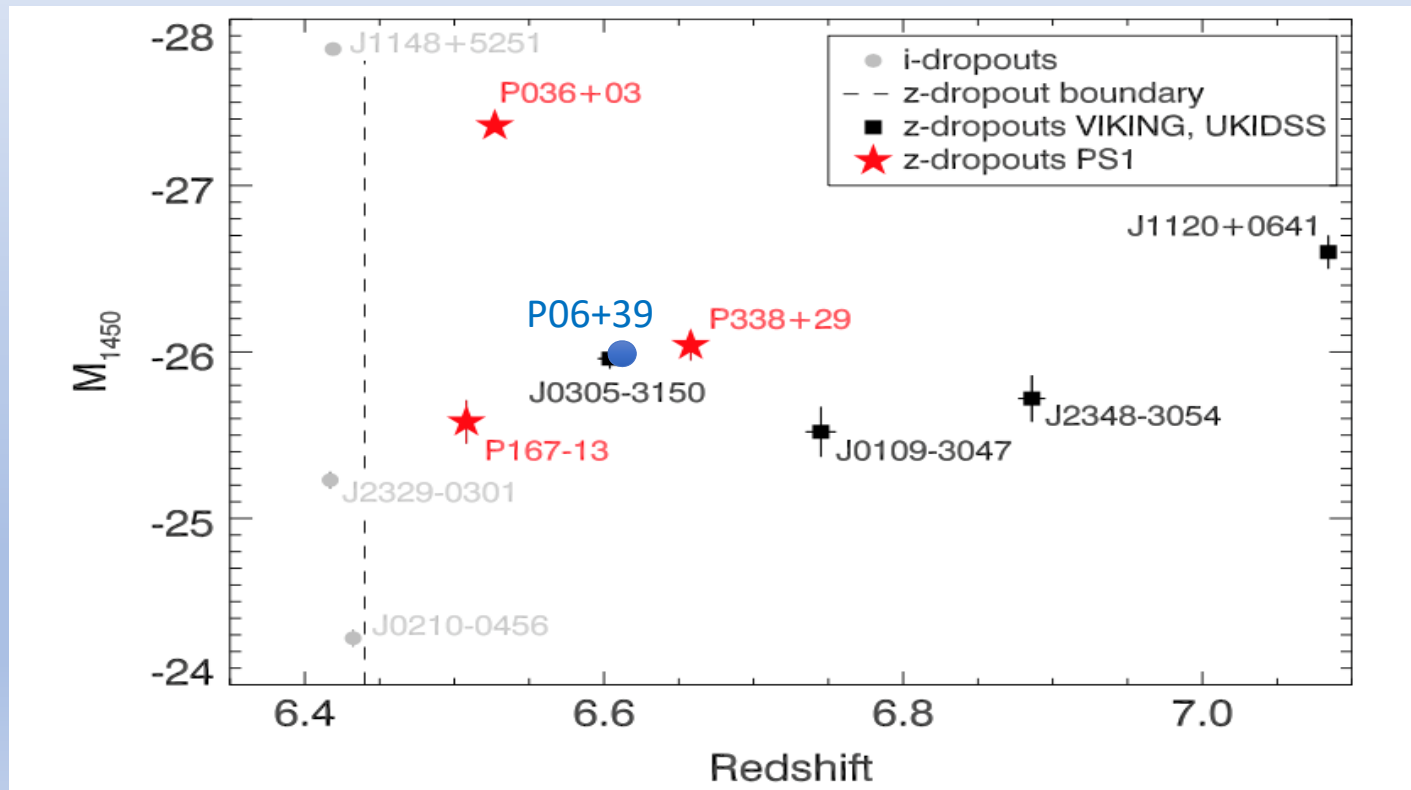
(Tang+17)

Quasar J006+39

- Redshift: 6.61 ± 0.02
- 7th farthest quasars known so far
- Number of quasars discovered above $z > 6.5$ (z-dropout): nine

Quasar J006+39

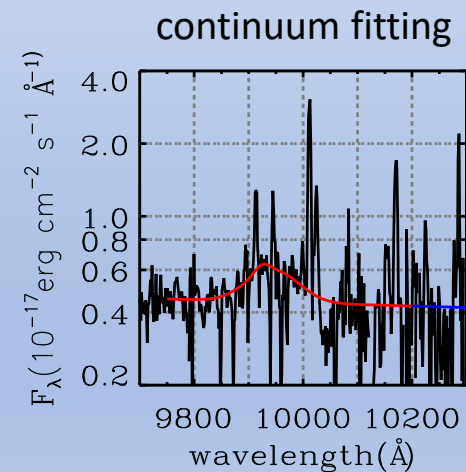
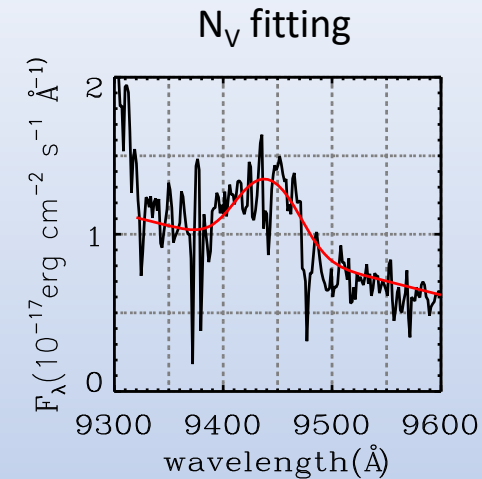
- Compare to other z-dropout Quasars



(Venemans+15)

Quasar J006+39 Fitting

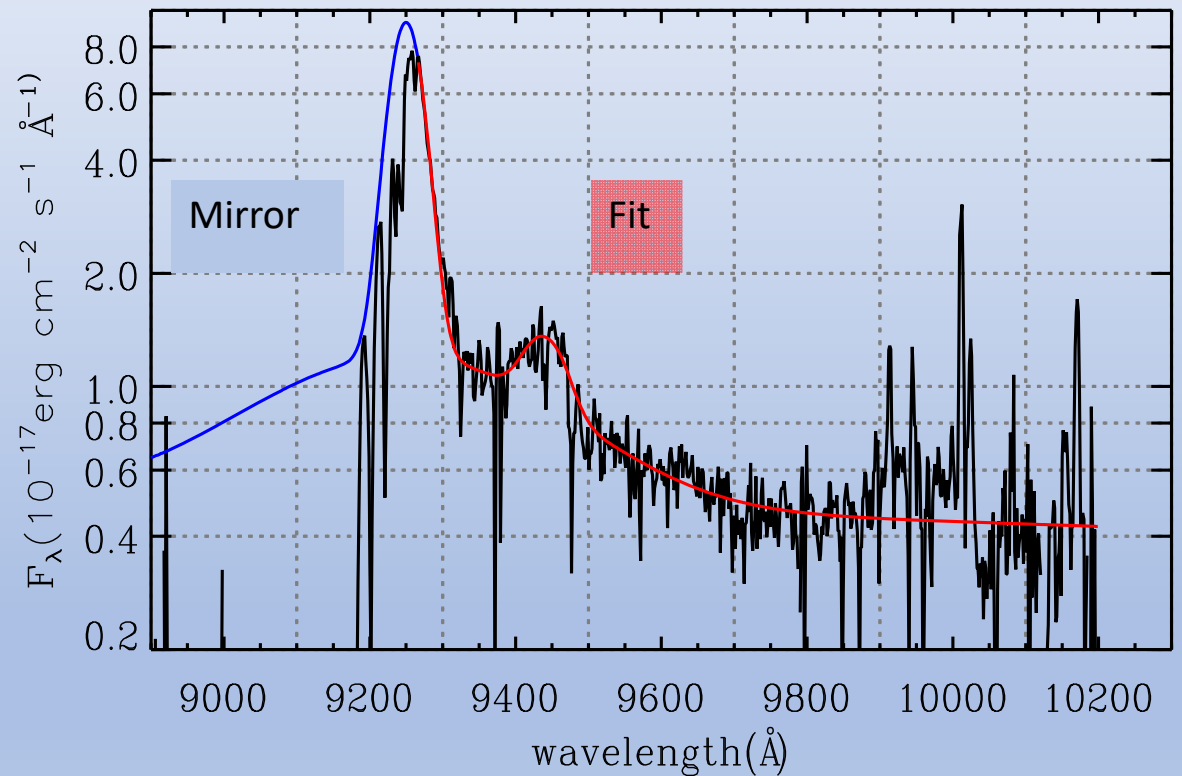
- NV fitting
 - Double Gaussian + linear continuum
 - Redshift measurement
- Continuum fitting
 - Double Gaussian (OI+Sill) + power-law
 - Luminosity estimation
 - Extrapolate to redshifted 1450Å



Quasar J006+39 Fitting

- Ly-alpha fitting
 - Compensate for absorption
 - Transmission measurement

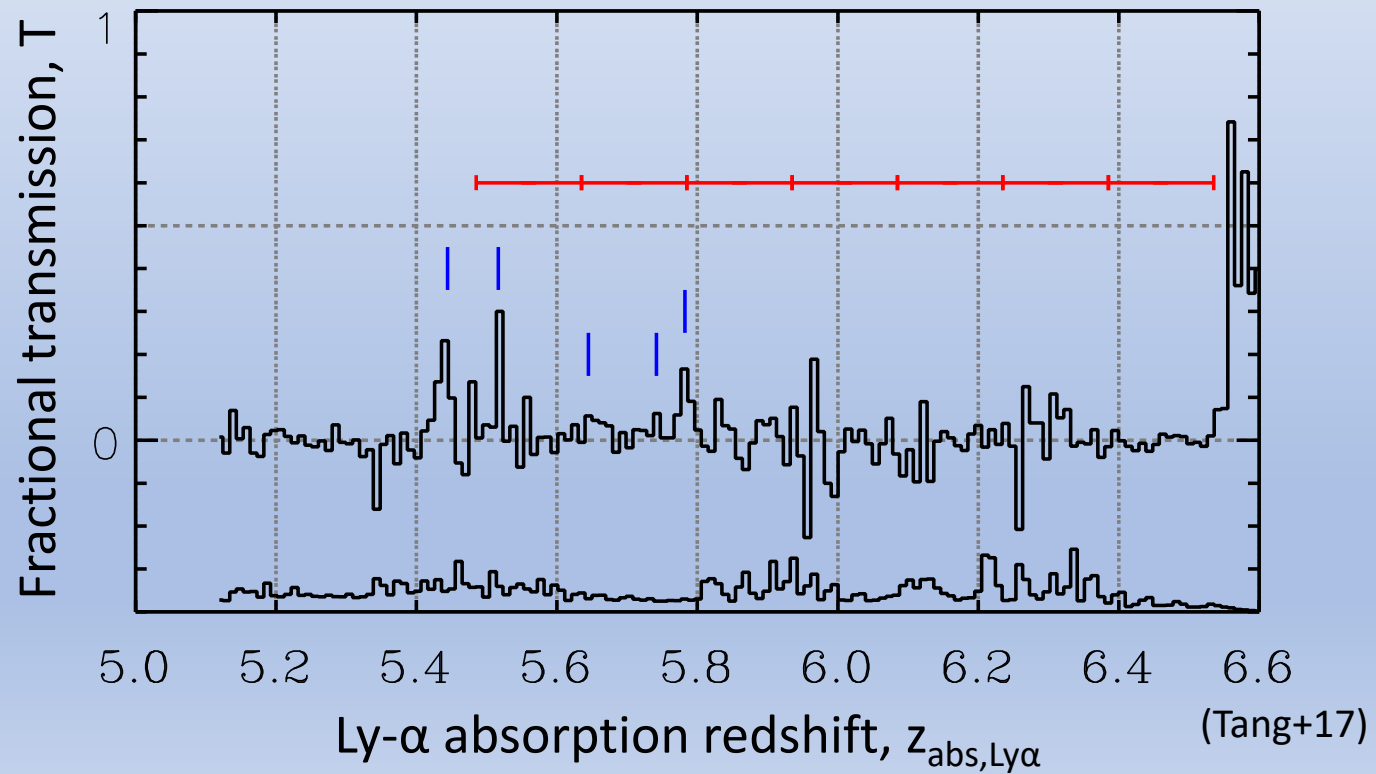
$$\text{Transmission} = \frac{\text{Observed spectrum}}{\text{Fitted spectrum}}$$



(Tang+17)

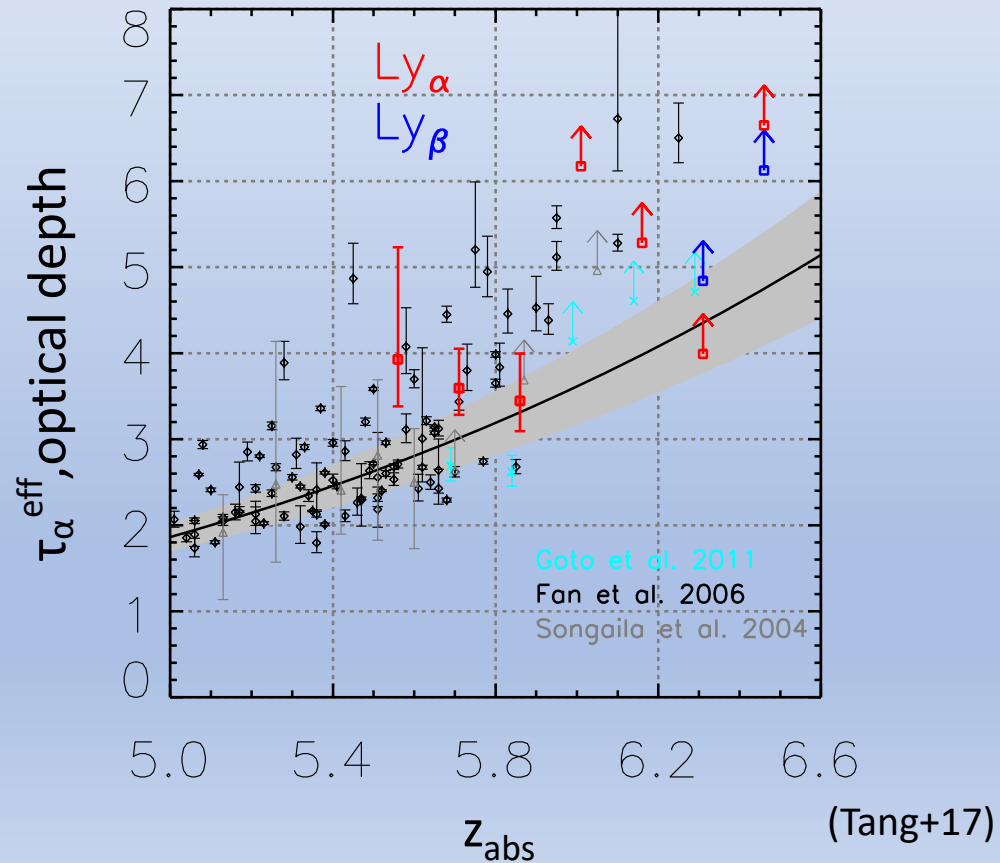
Transmission Measurement - Optical Depth

- Gunn-Peterson effective Optical Depth



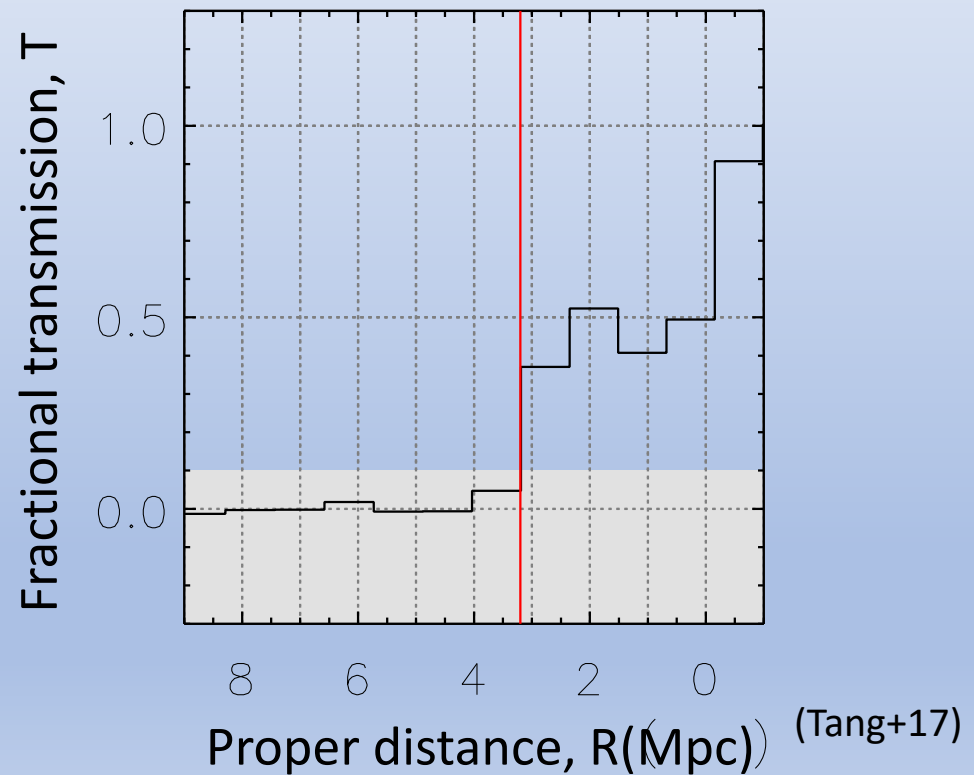
Transmission Measurement - Optical Depth

- Compare to previous optical depth measurement
 - Ly- α
 - Ly- β

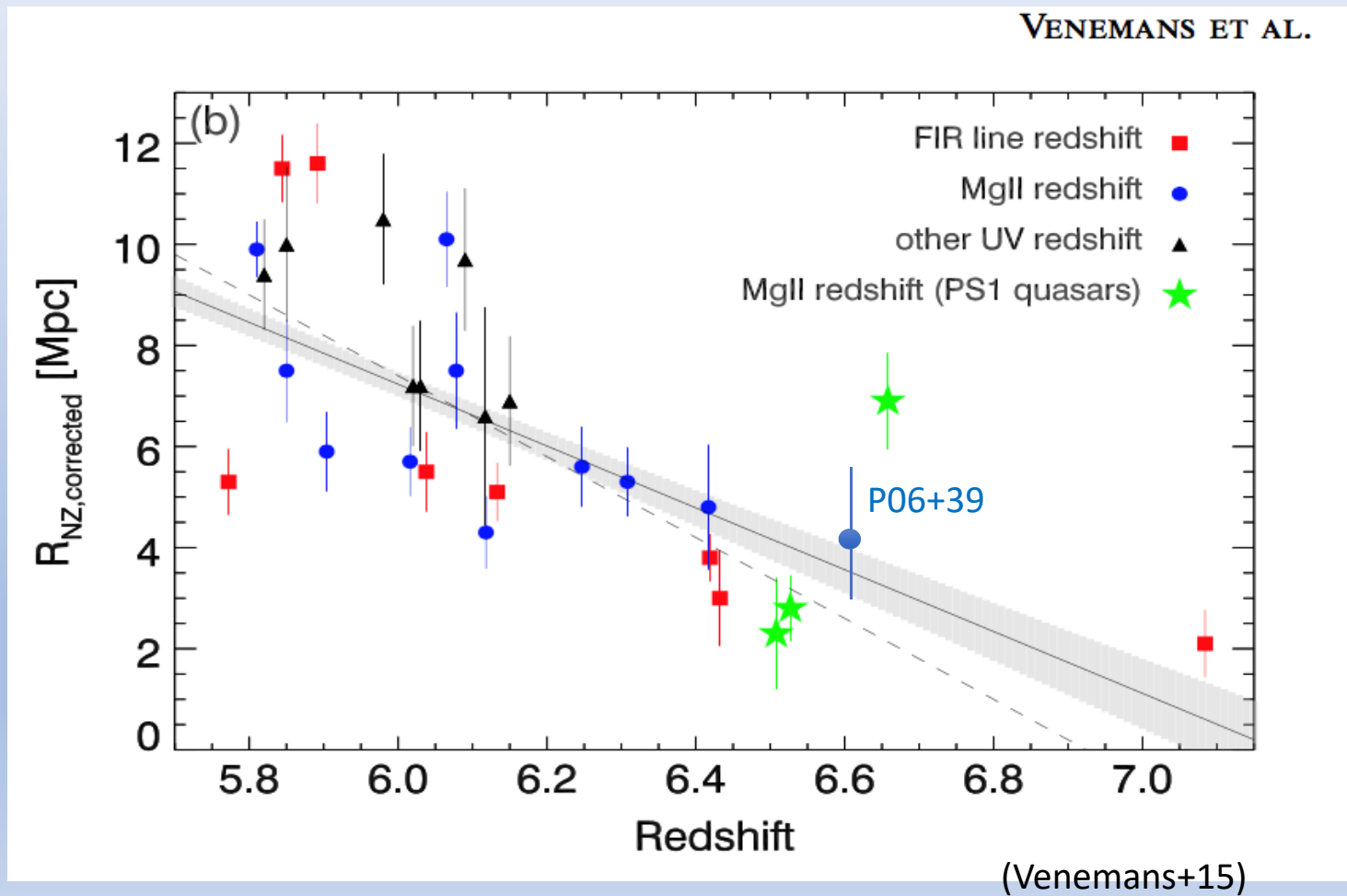


Transmission Measurement - Near Zone Region

- Near zone region:
 - $R_{p,NZ} = 3.2 \pm 1.1\text{Mpc}$
- Corrected near zone region:
 - $R_{NZ,corr} = 4.3 \pm 1.5\text{Mpc}$
 - corrected for luminosity



Transmission Measurement - Near Zone Region



Summary

- Found a QSO at Redshift 6.61 (7th farthest quasars known so far)
- Gunn-Peterson optical depth measurement provides more data of neutral hydrogen fraction at high-z universe
- Near-zone region follows the trend: Ionization ability is similar to other high-z quasars

Transmission Measurement - Near Zone Region

- Near-zone Region
 - Transmission $T < 0.1 (z_{GP})$
 - $R_{p,NZ} = (D_Q - D_{GP}) / (1 + z_Q)$
 - The D_Q and D_{GP} are the comoving distances of the redshift of the quasar host galaxy (z_Q) and the redshift where the transmission drops below 0.1 (z_{GP}), respectively.
 - $3.2 \pm 1.1 \text{Mpc}$
- Corrected Near-zone region
 - Consider Luminosity
 - $R_{NZ,corr} = 10^{0.4 \times (27 + M_{1450,AB}) / 3} R_{NZ}$
 - $4.3 \pm 1.5 \text{Mpc}$

Transmission Measurement - Optical Depth

- Convert Transmission into effective optical depth:

$$\tau_{eff} = -\ln(T)$$

- Ly-alpha optical depth (τ_α)
 - $5.56 < z < 6.46$ (Avoid Ly-alpha & beta emission)
- Ly-beta optical depth (τ_β)
 - Estimate Ly-alpha absorption:
 - $\tau_\alpha = (0.85 \pm 0.06) \times \left(\frac{1+z}{5}\right)^{(4.3 \pm 0.3)}$ from (Fan+06)
 - below 1017 \AA (Avoid Ly-beta emission)
- $\tau_\alpha / \tau_\beta = 2.25$