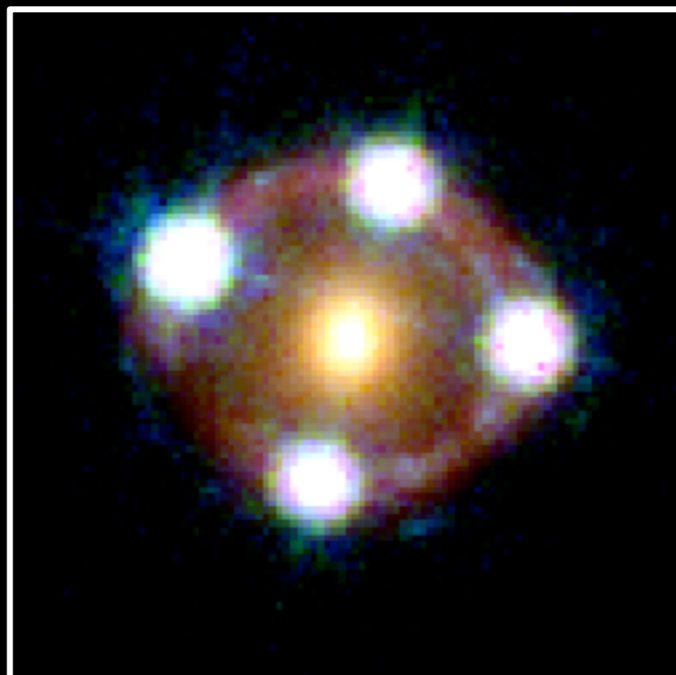
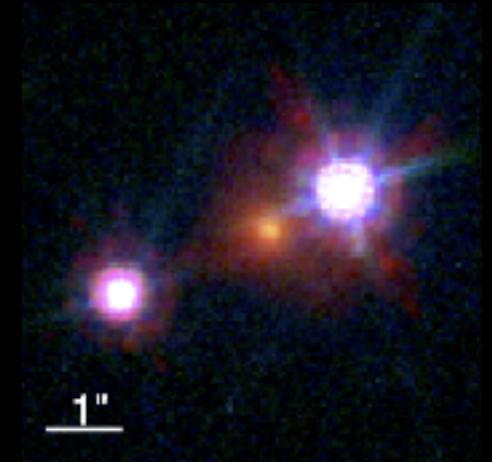
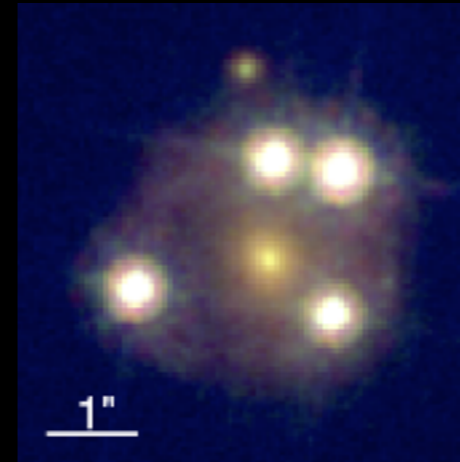
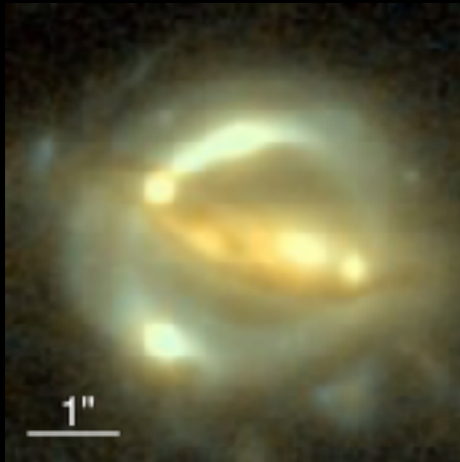


Cosmology from Gravitational Lens Time Delays

Updates on the H0LiCOW Project



Kenneth Wong

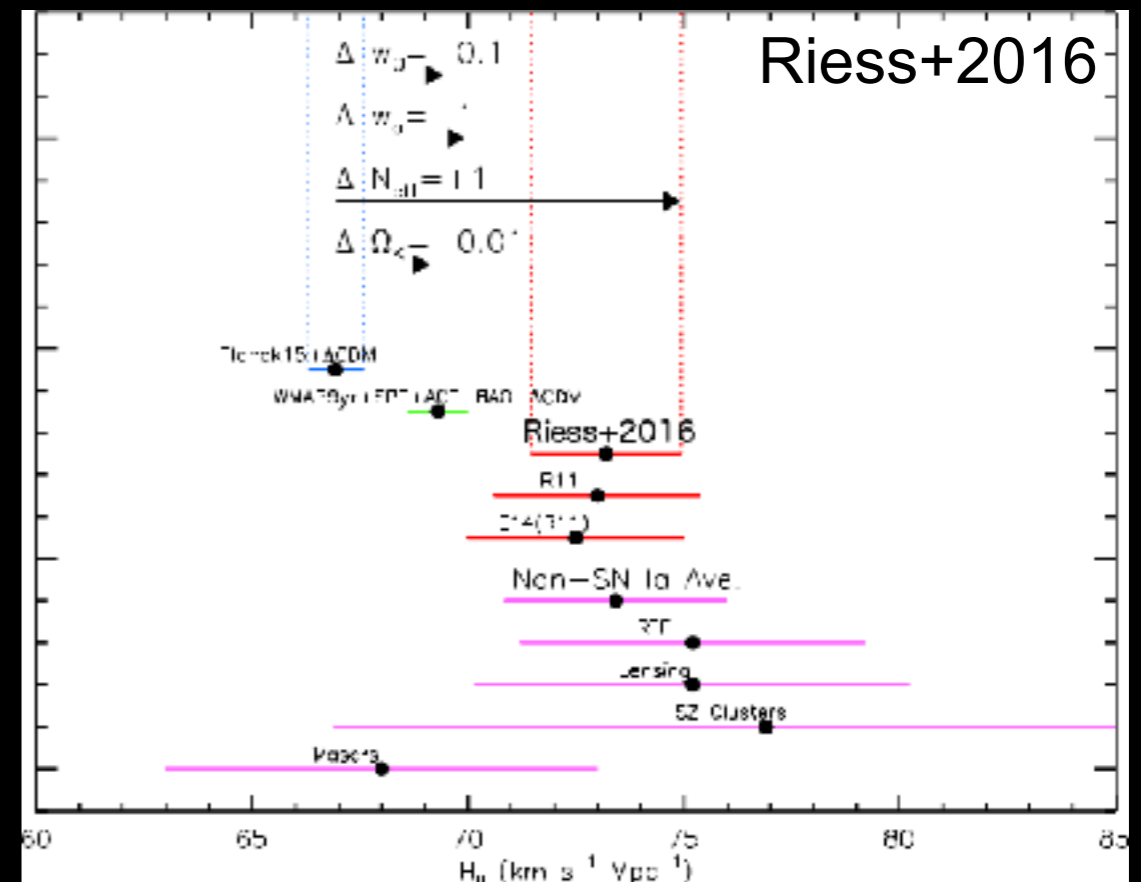
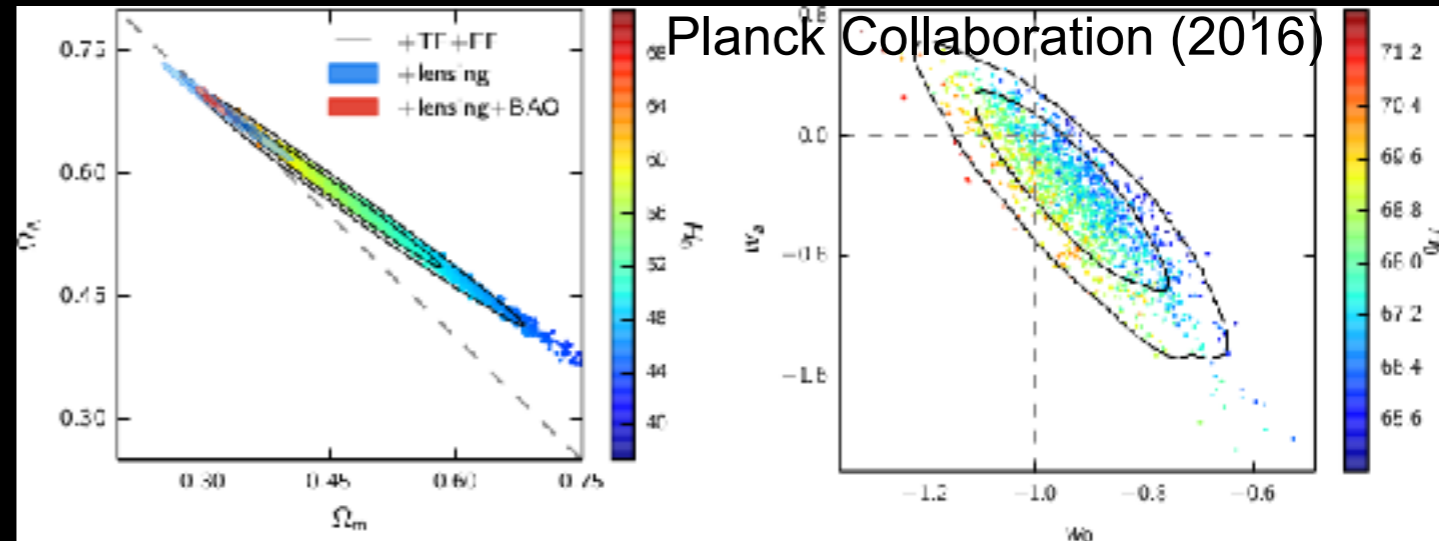
EACOA Fellow

National Astronomical Observatory of Japan

14th Asia-Pacific Regional IAU Meeting
July 4, 2017

Independent H_0 Measurements

- The Hubble Constant (H_0) sets the expansion rate of the universe
- *Planck* flat Λ CDM results suggest an H_0 value lower than other measurements
- Independent distance ladder results (Riess+2016) favor a higher H_0
- Tension? New physics? Need more precise and accurate measurement of H_0



Strong Gravitational Lensing

- Background object (source) magnified by foreground object (lens)
- Multiple images → create lens model
- Can determine mass distribution of lens if mass along line of sight is accounted for
- Can be used to constrain cosmology

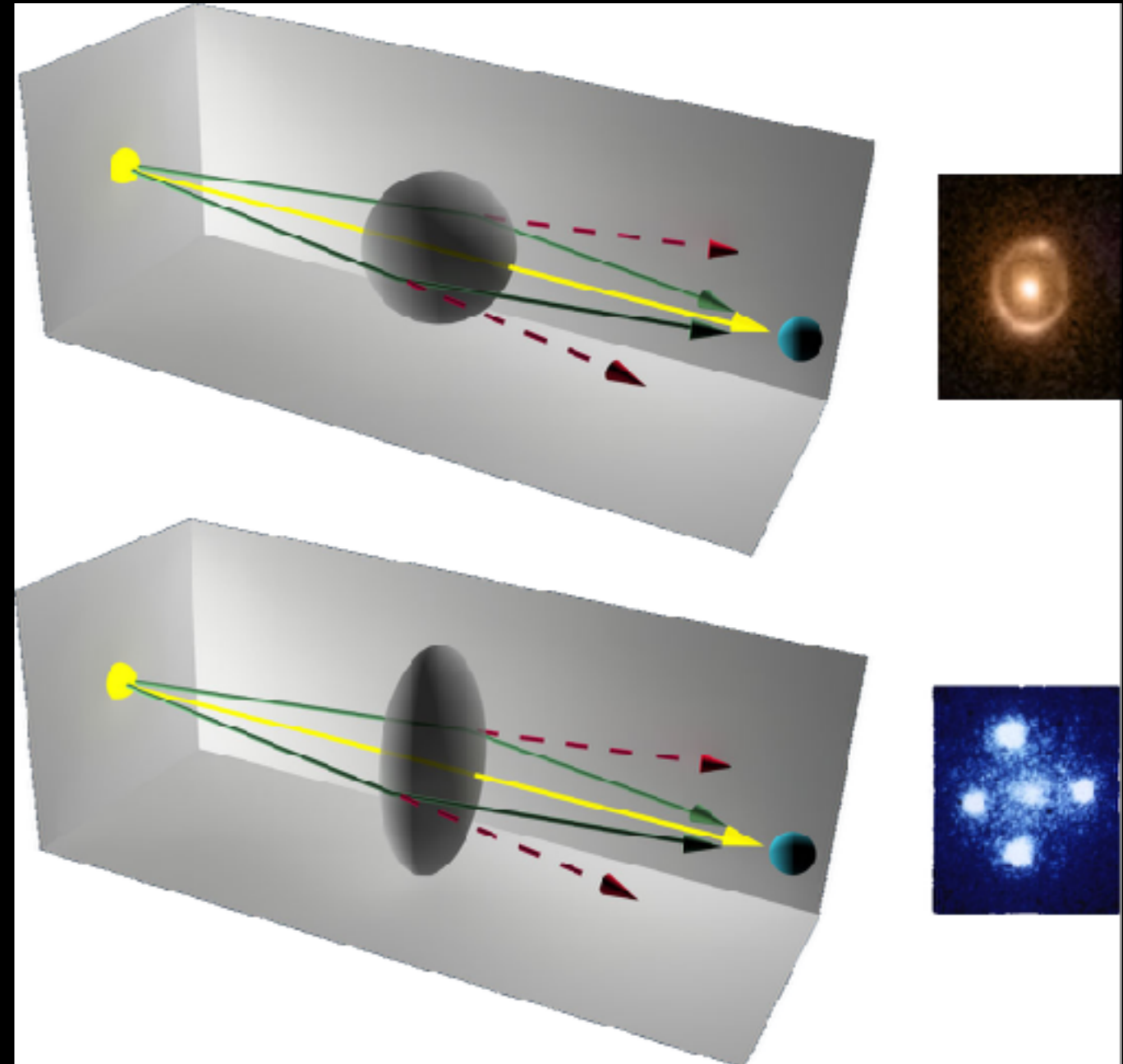
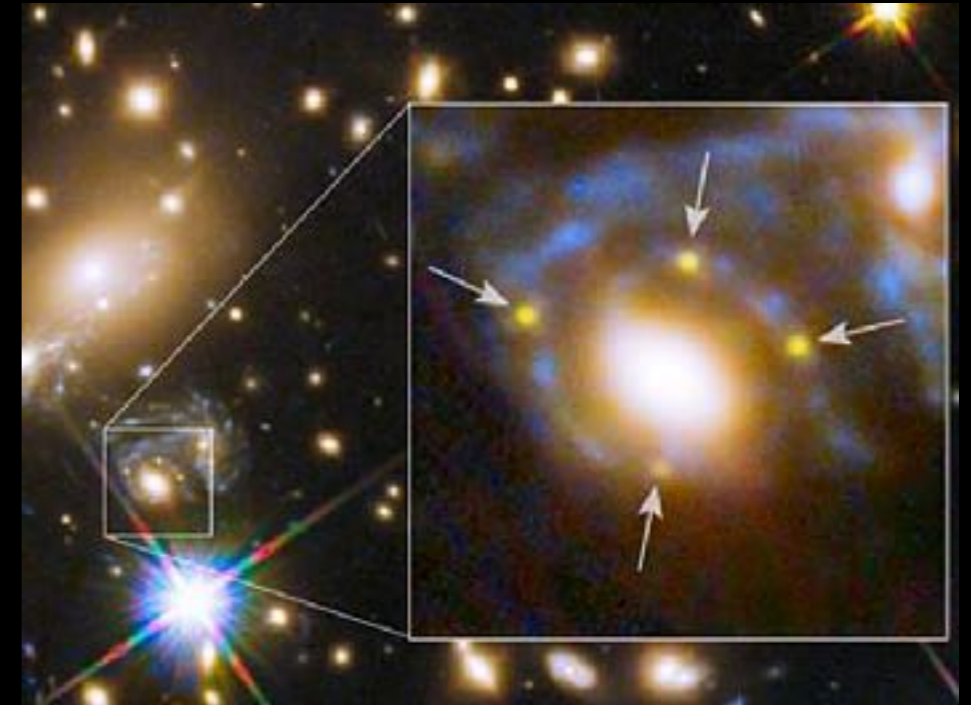


Image credit: ASIAA EPO

Gravitational Lensing Time Delays

- If source is variable, there is a “time delay” between the multiple images
- Can determine “time-delay distance”, inversely proportional to H_0
- One-step method to infer H_0 , independent of distance ladder
- e.g., lensed supernovae (e.g. Kelly+2015, Goobar+2016), but very rare!



SNe “Refsdal”

Time delay Lens potential (from mass model)

$$\Delta t = \frac{1}{c} D_{\Delta t} \phi_{lens}$$

Time-delay distance

$$D_{\Delta t} \propto H_0^{-1}$$

Gravitational Lensing Time Delays

- Lensed quasars
 - variable on short timescales (~days)
 - can be monitored to measure time delay
 - bright and easy to detect

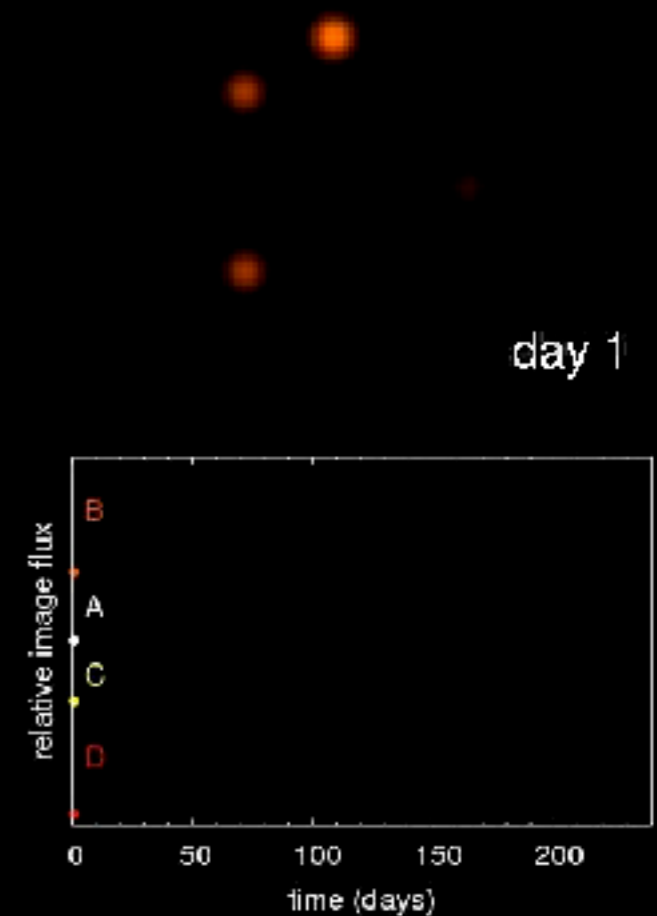
Time delay

Lens potential (from mass model)

$$\Delta t = \frac{1}{c} D_{\Delta t} \phi_{lens}$$

Time-delay distance

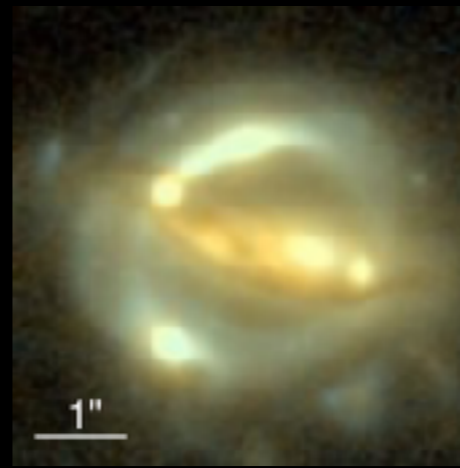
$$D_{\Delta t} \propto H_0^{-1}$$





H0LiCOW: H_0 Lenses in COSMOGRAIL's Wellspring

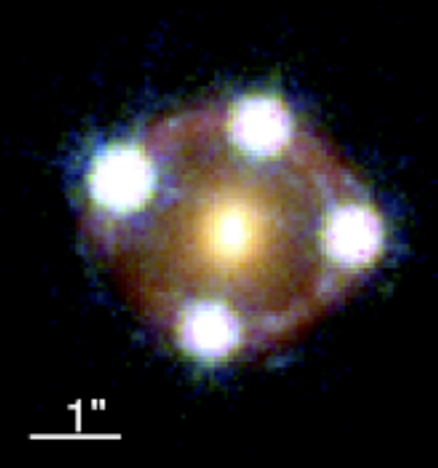
- Detailed analysis of five time-delay lenses (Suyu+2017)
 - long term monitoring from COSMOGRAIL
 - high-resolution *HST* imaging for detailed lens modeling
 - imaging/spectroscopy to characterize mass along line of sight
- Will constrain H_0 to $< 3.5\%$ precision
- First two lenses previously analyzed (Suyu+2010, 2013)
- Four additional lenses tentatively will be added to sample ($\sim 2\%$ precision on H_0)



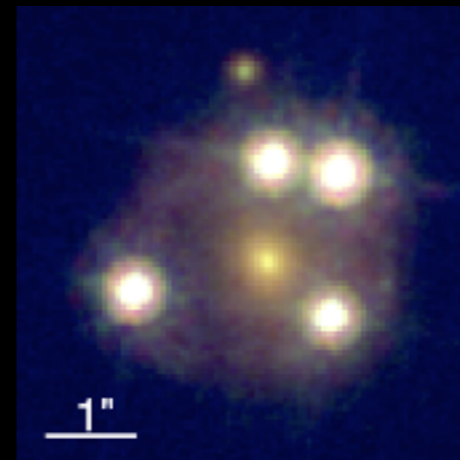
B1608+656



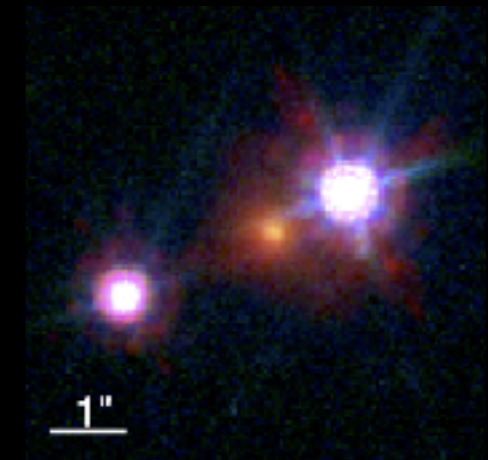
RXJ1131-1231



HE 0435-1223



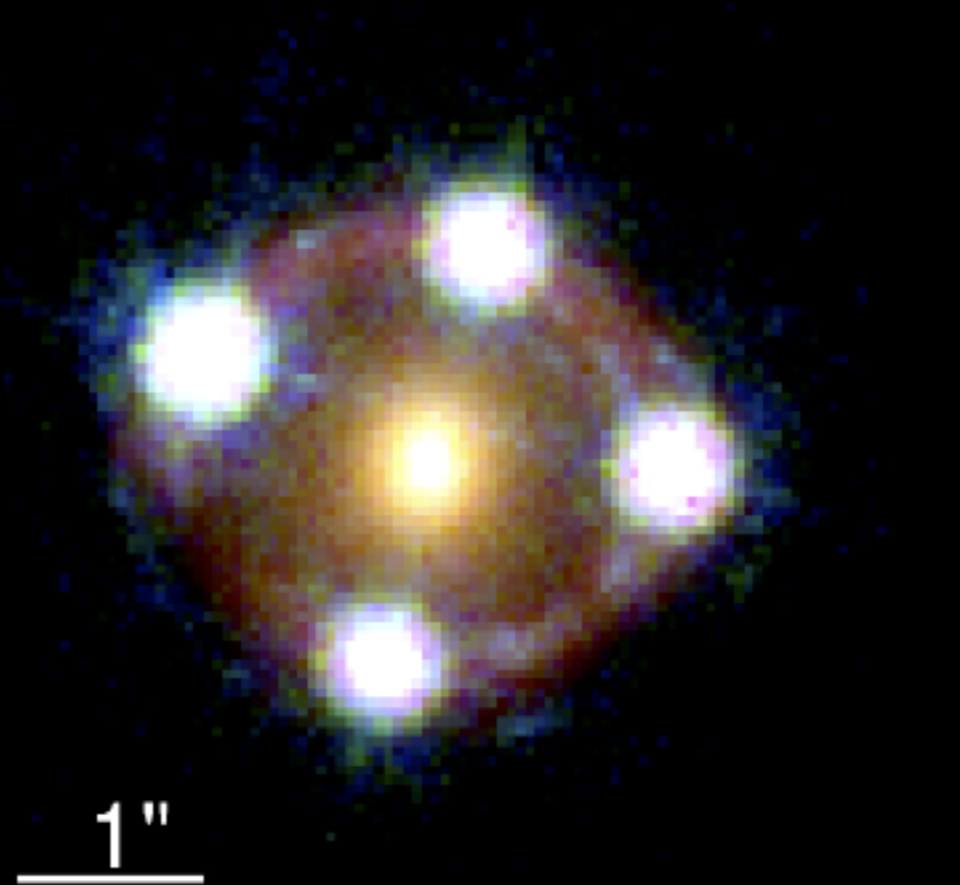
WFI2033-4723



HE 1104-1805

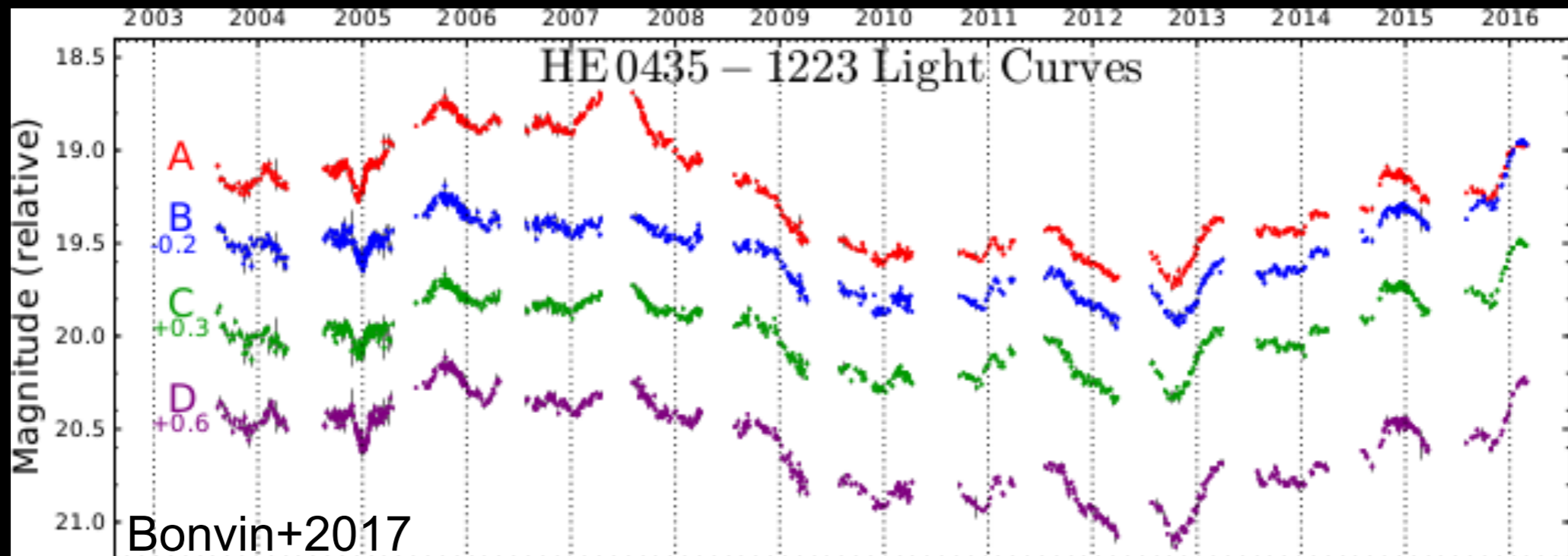
Latest Results from HE0435

- Extensive dataset
 - *HST* imaging in 3 bands (F555W, F814W, F160W)
 - 13-year monitoring by COSMOGRAIL for accurate time delays
 - Lens velocity dispersion from Keck/LRIS
 - Spectroscopic data on LOS galaxies to get perturber redshifts
 - Multiband photometry to get photo-zs and stellar masses of LOS galaxies
- Developed new PSF reconstruction and multi-plane lensing techniques (Suyu, KW+ in prep.)
- Full analysis and results:
 - Sluse+2017 (LOS galaxy spectroscopy)
 - Rusu+2017 (LOS photo-zs/stellar masses)
 - KW+2017 (Lens model)
 - Bonvin+2017 (Time-delay measurements)



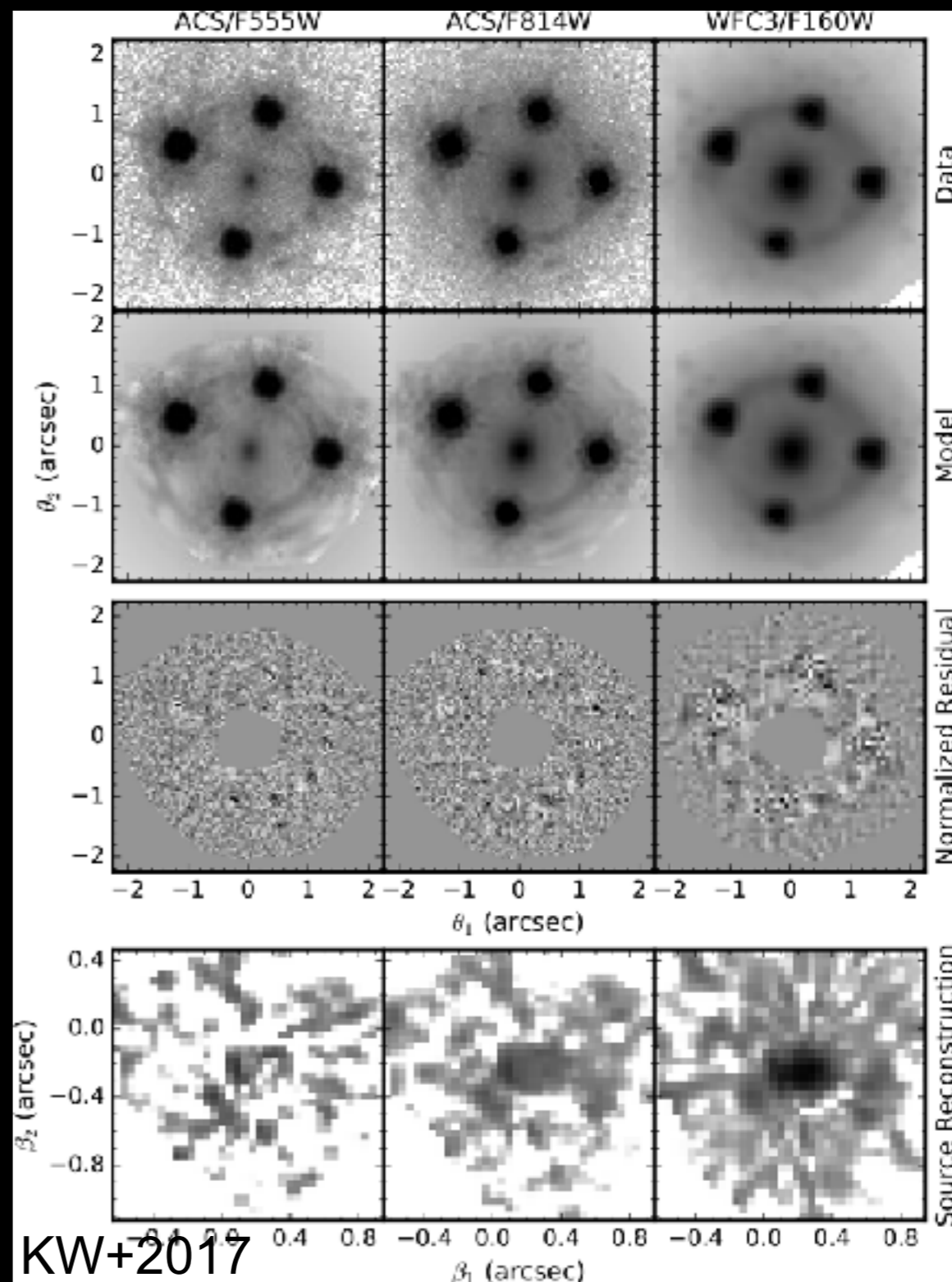
Accurate Time Delays

- COSMOGRAIL: long-term monitoring of time-delay lenses using small (1-m and 2-m) telescopes (Courbin+2011)
- Well-tested algorithms for time-delay measurements (Tewes+2013)
- Time delays of HE0435 from 13 years of monitoring (Bonvin+2017)
 - Long time baselines needed to minimize effects of microlensing

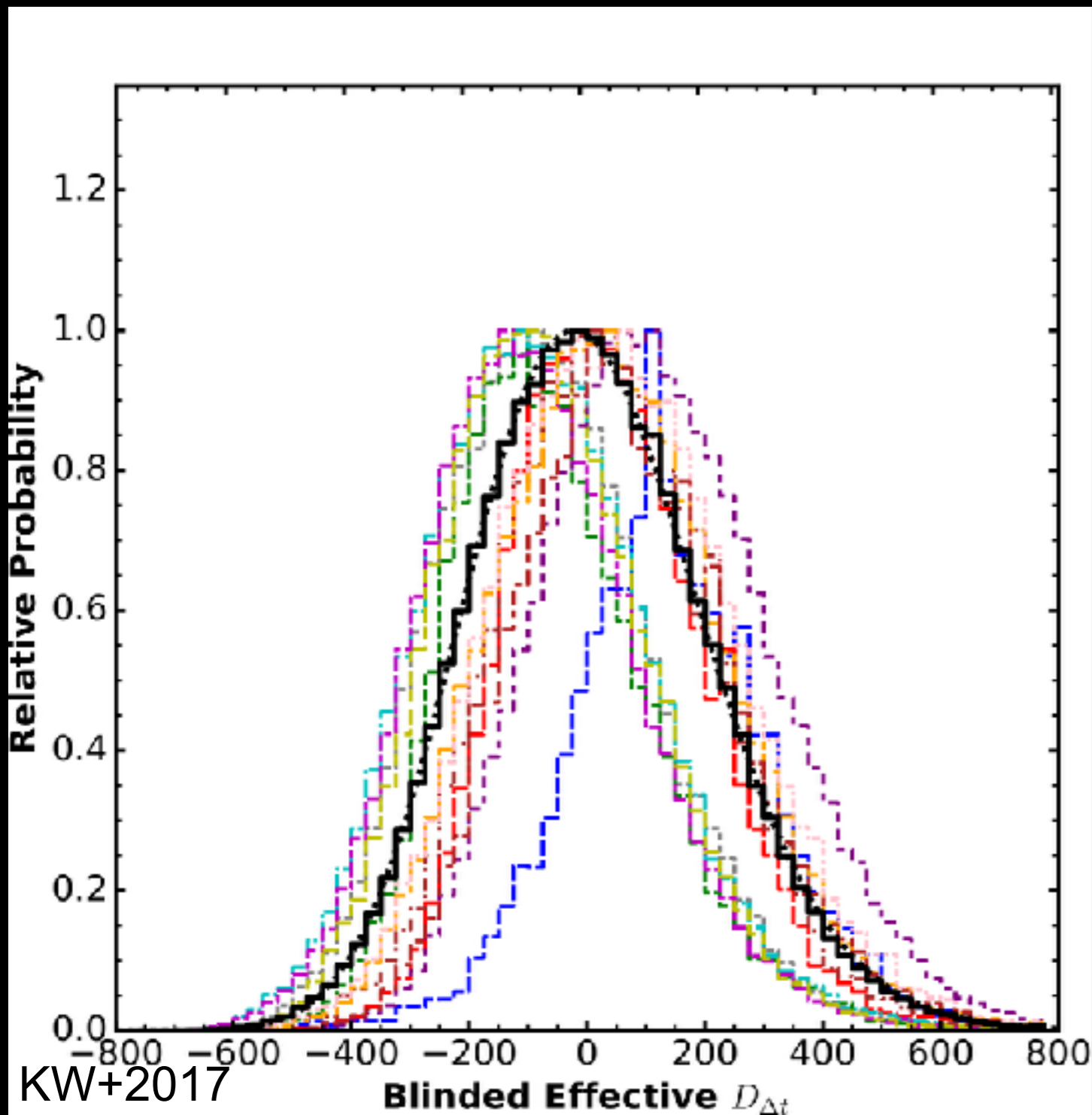


Accurate Lens Model

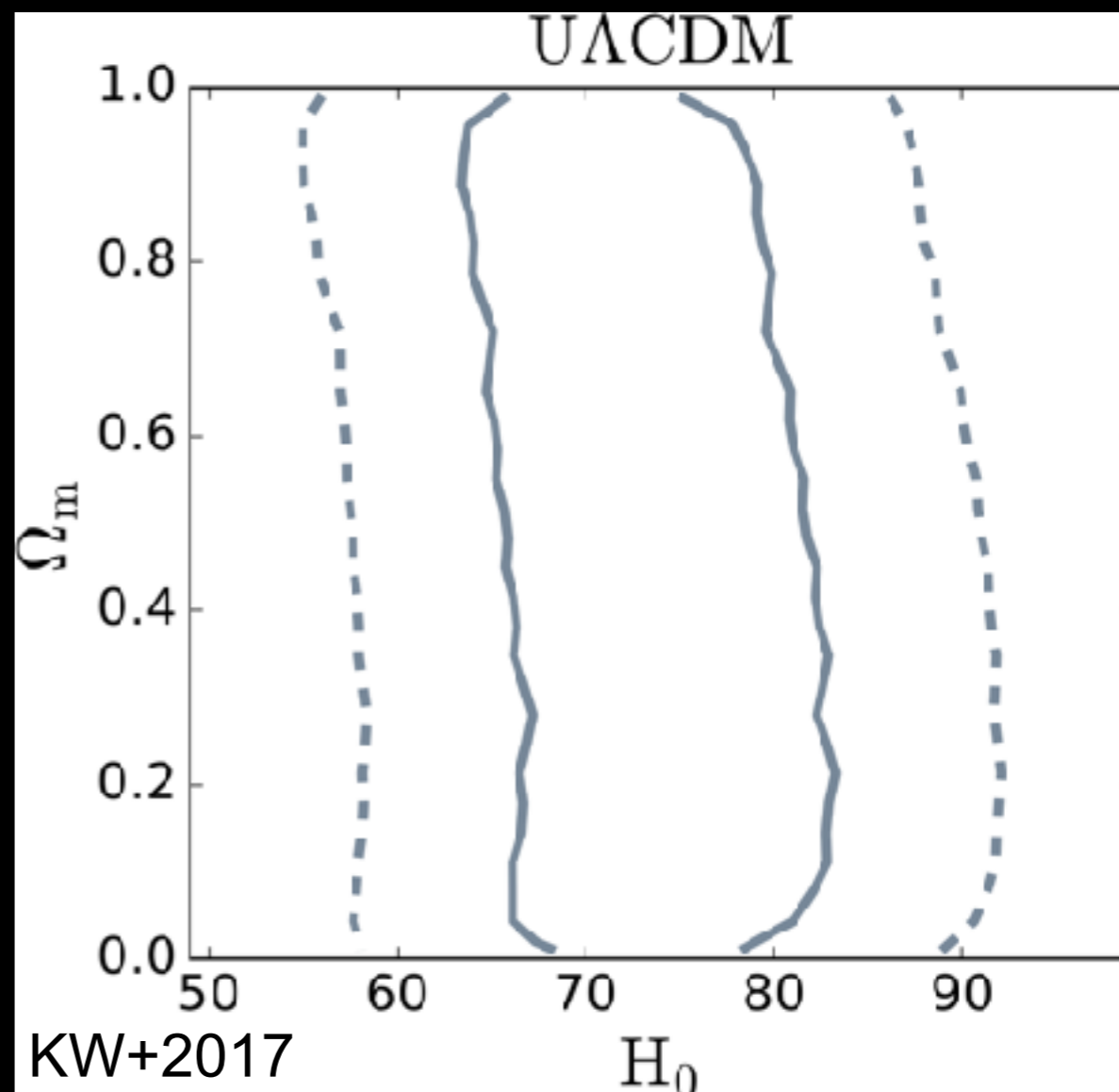
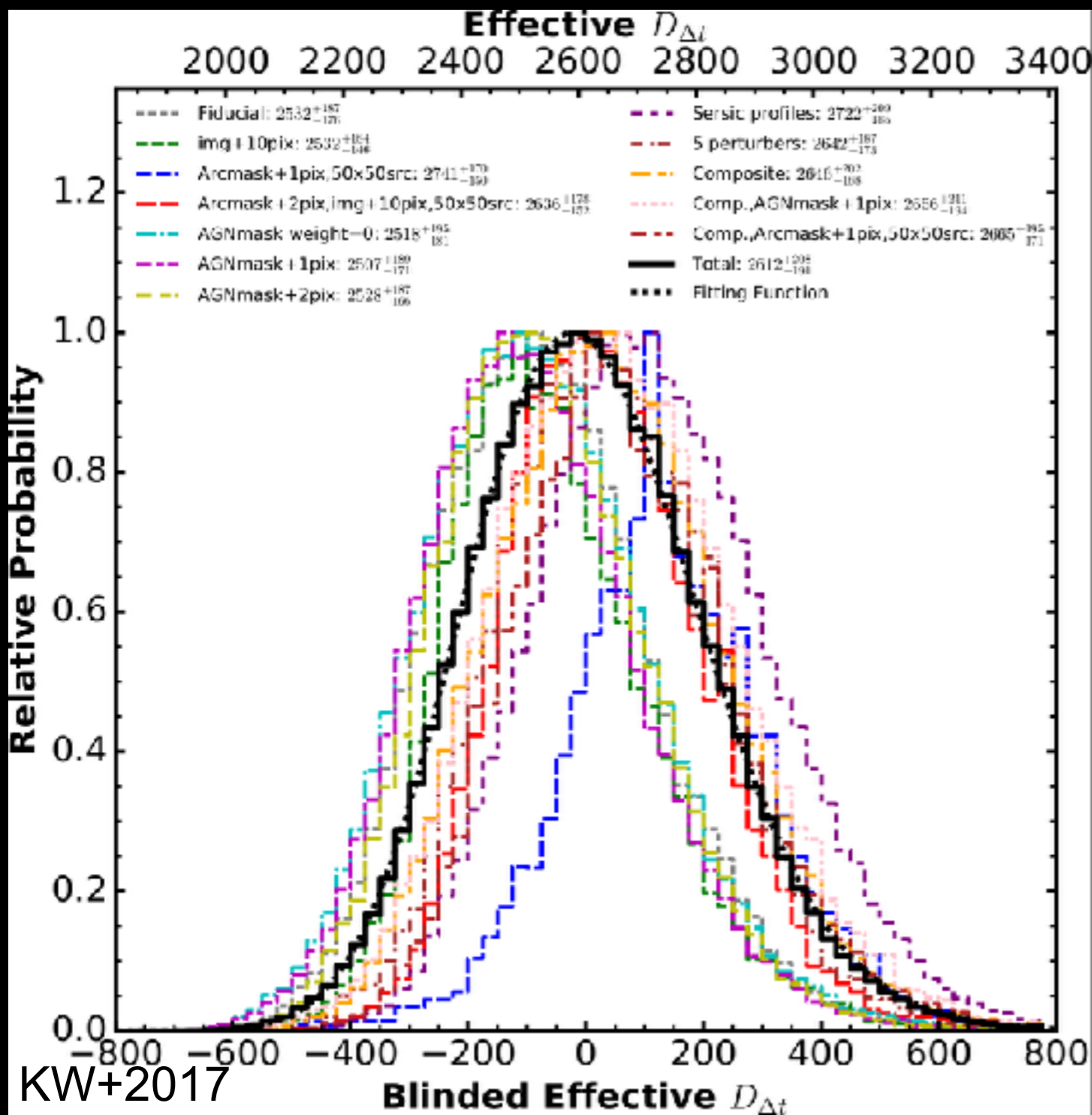
- Accurate lens model using 3-band of *HST* imaging (KW+2017)
- High-resolution needed to model quasar host galaxy
- Adaptive PSF correction using quasar images (e.g. Chen+2016)
- Account for nearby perturber using multi-plane lensing formalism
- External convergence from weighted galaxy number counts
- Velocity dispersion of lens galaxy from Keck/LRIS spectrum to break model degeneracies



Blind Analysis

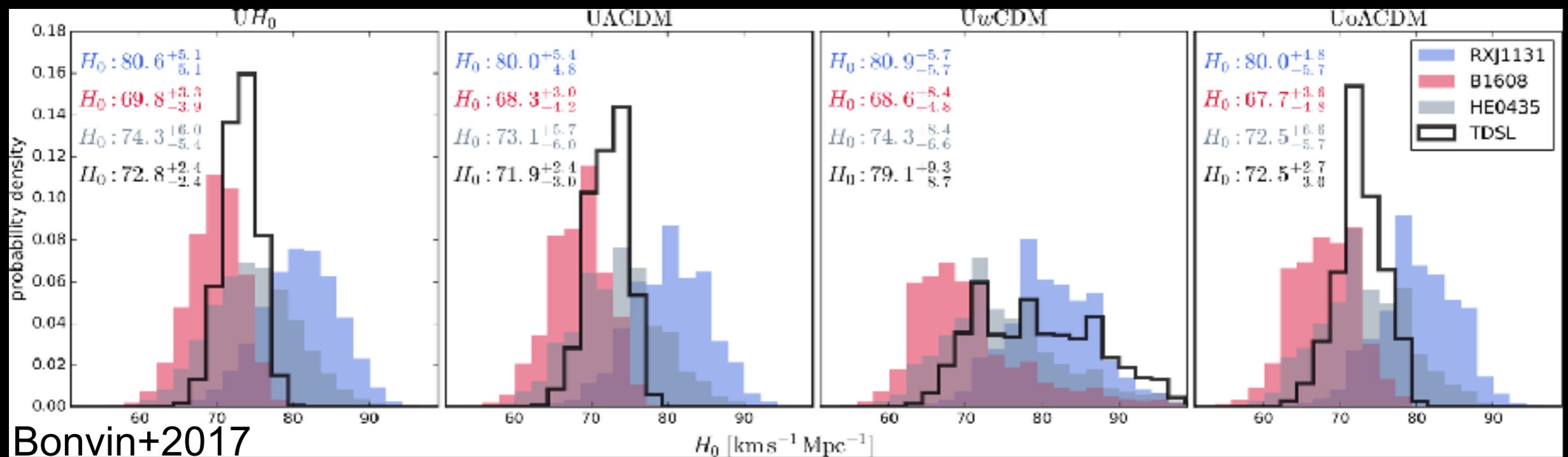


Results from HE0435



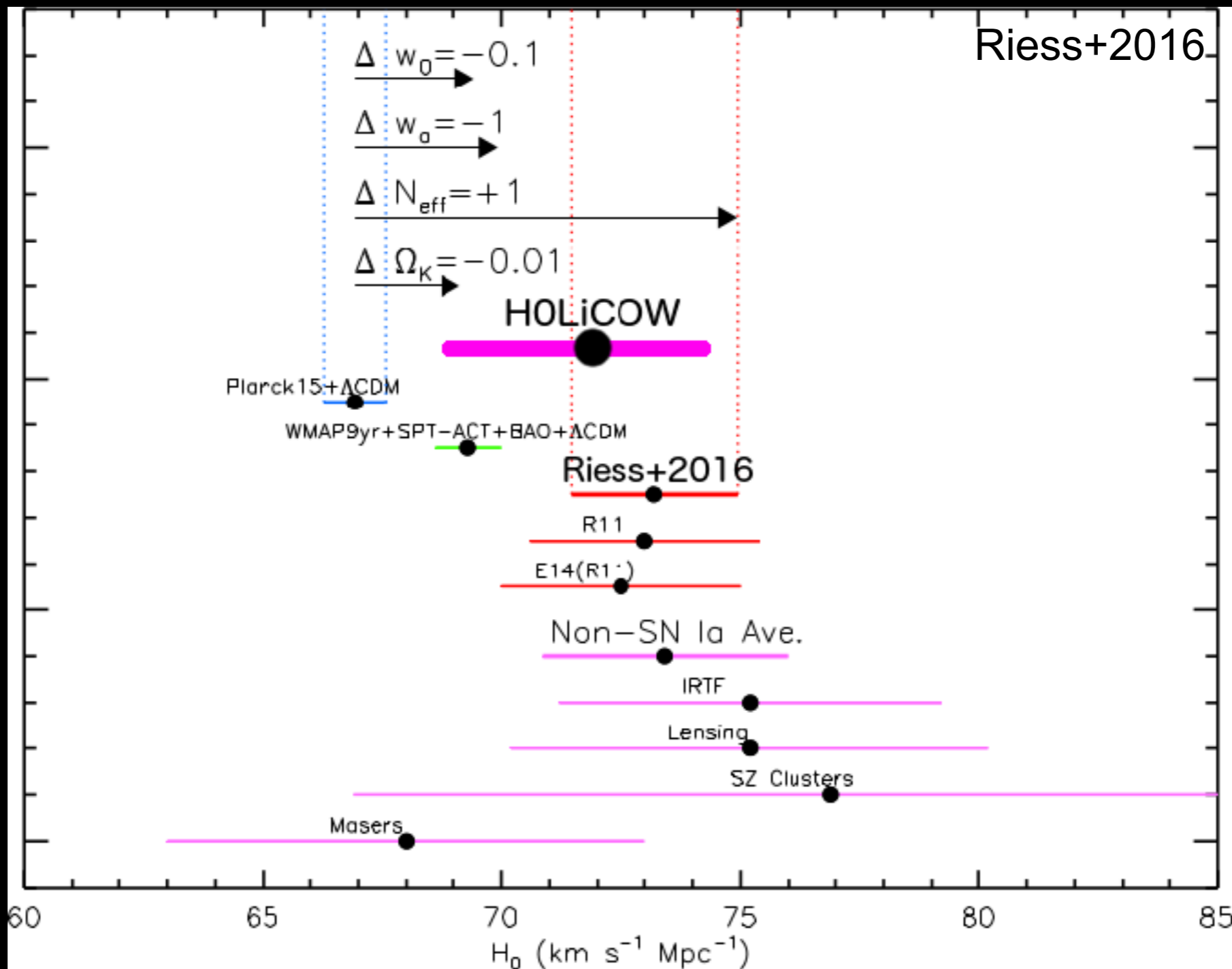
~8% constraint on $D_{\Delta t}$
 $H_0 = 73.1^{+5.7}_{-6.0}$ km/s/Mpc
 for flat Λ CDM cosmology

Combined Results from 3 H0LiCOW Lenses

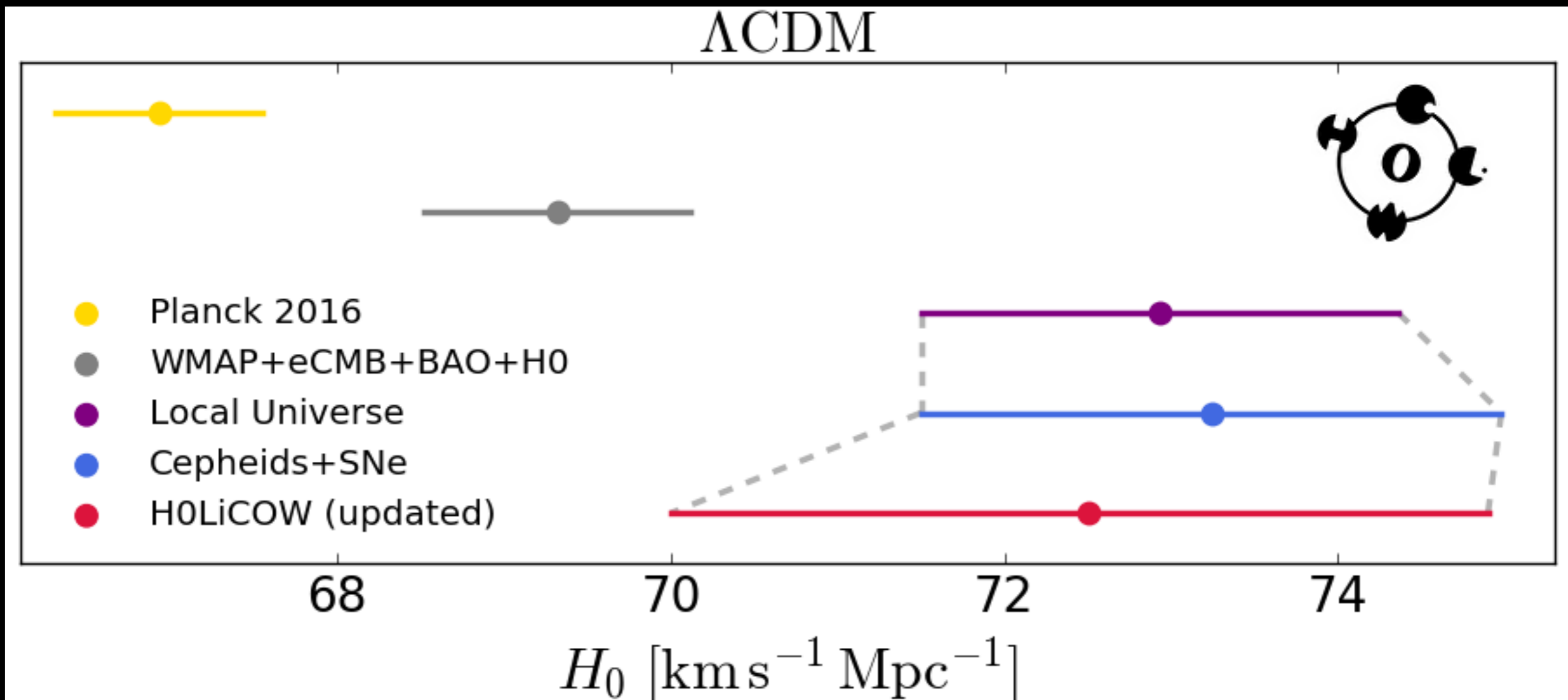


~3.8% precision on H_0 from 3 H0LiCOW lenses
 $H_0 = 71.9^{+2.4}_{-3.0}$ km/s/Mpc for flat ΛCDM cosmology

Combined Results from 3 H0LiCOW Lenses

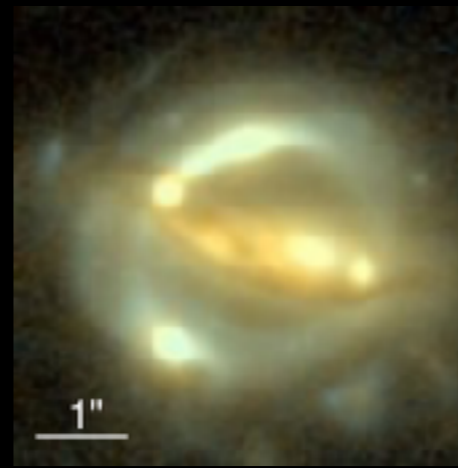


Combined Results from 3 H0LiCOW Lenses

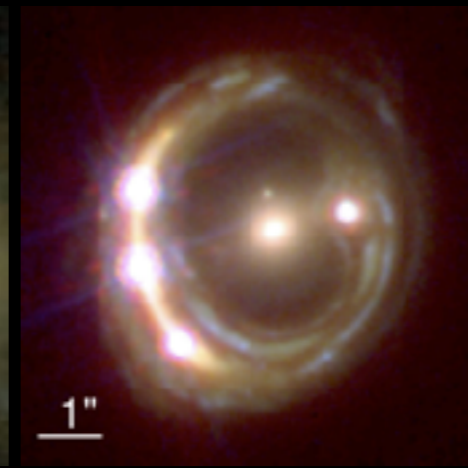


Ongoing Work

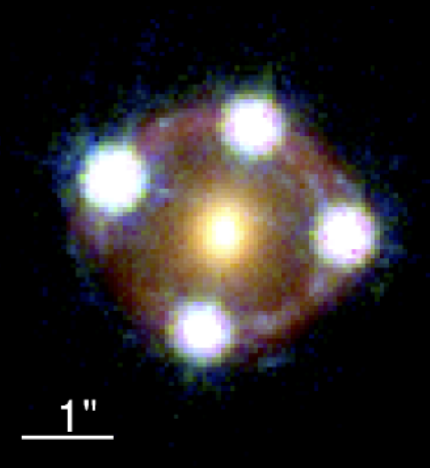
- Analysis of WFI2033 in progress, HE1104 to follow
- Four additional lenses with time delays and *HST* data
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 - full sample of 9 lenses will constrain H_0 to $\sim 2\%$
- Weak lensing analysis of HE0435 field (Tihhonova+ in prep.)
 - independent test of LOS mass distribution
- AGN host galaxy studies
 - study evolution of BH-bulge relation using lensing magnification (Ding+2017a)
 - application to H0LiCOW data (Ding+ 2017b)



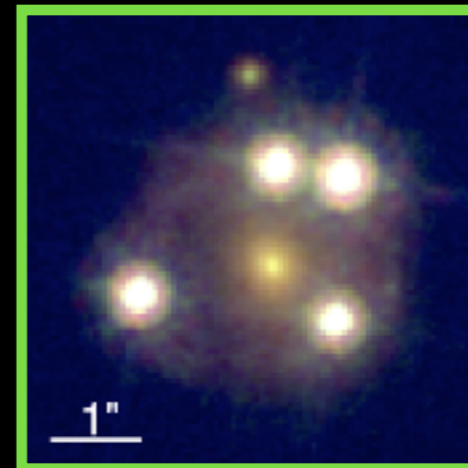
B1608+656



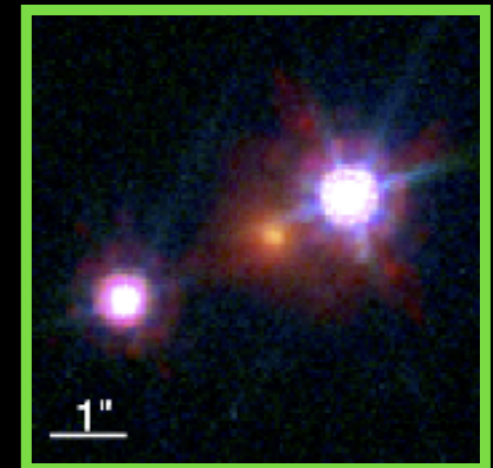
RXJ1131-1231



HE 0435-1223



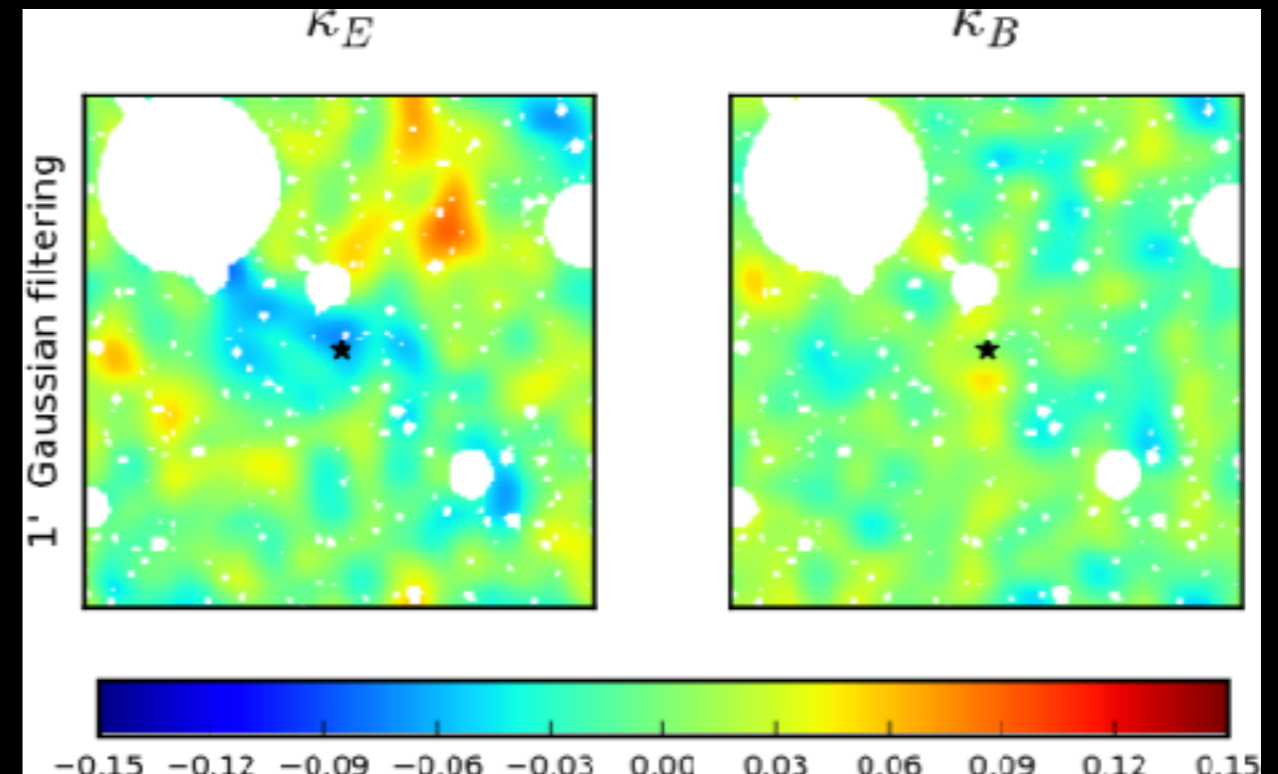
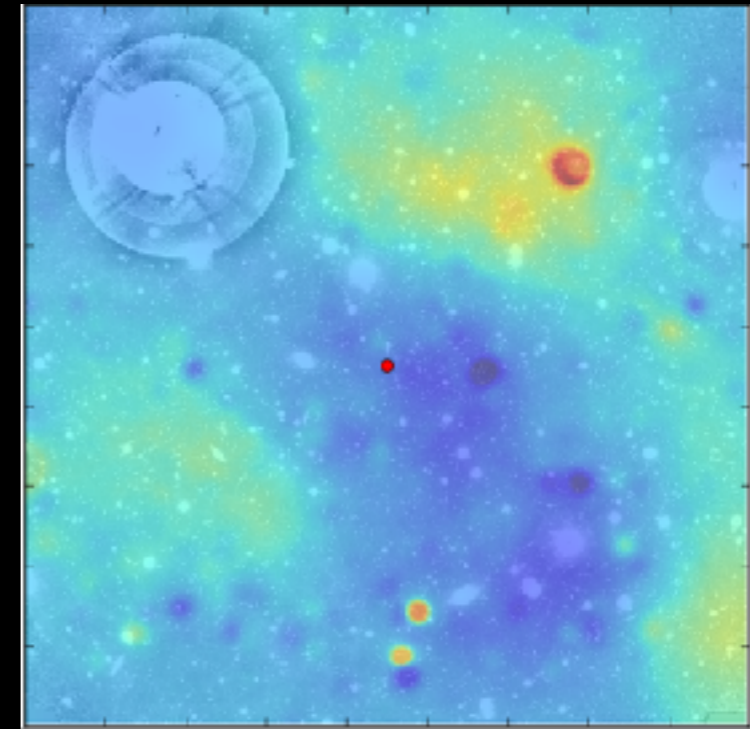
WFI2033-4723



HE 1104-1805

Ongoing Work

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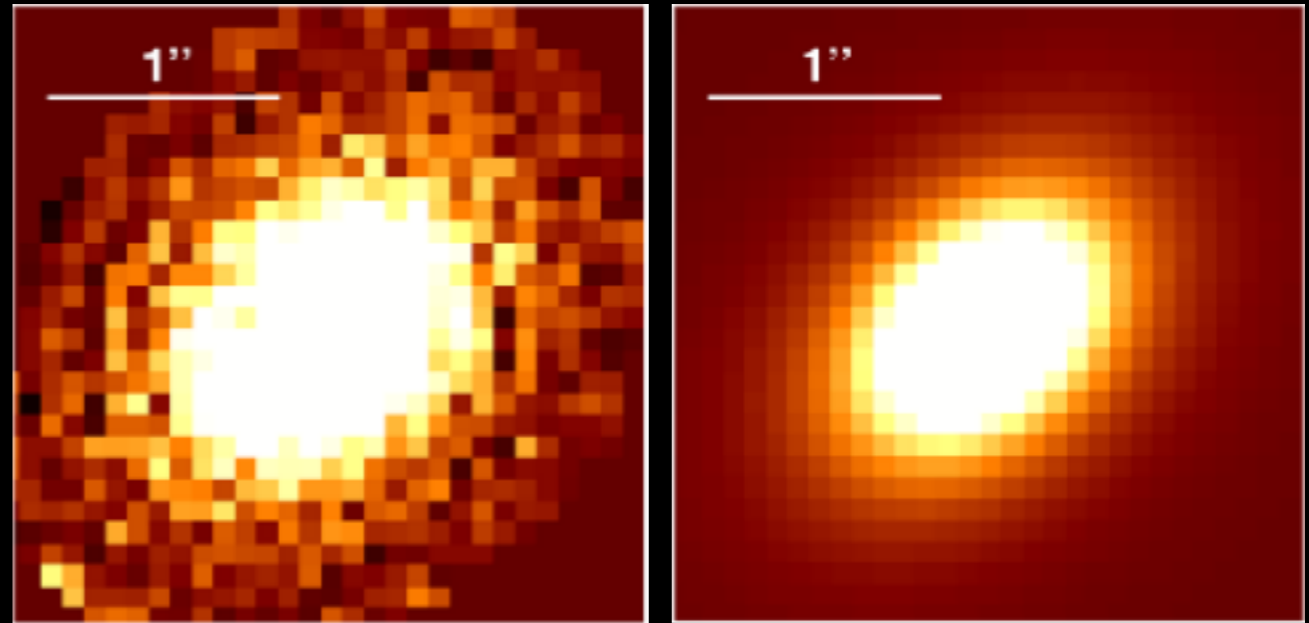


Tihhonova+ in prep.

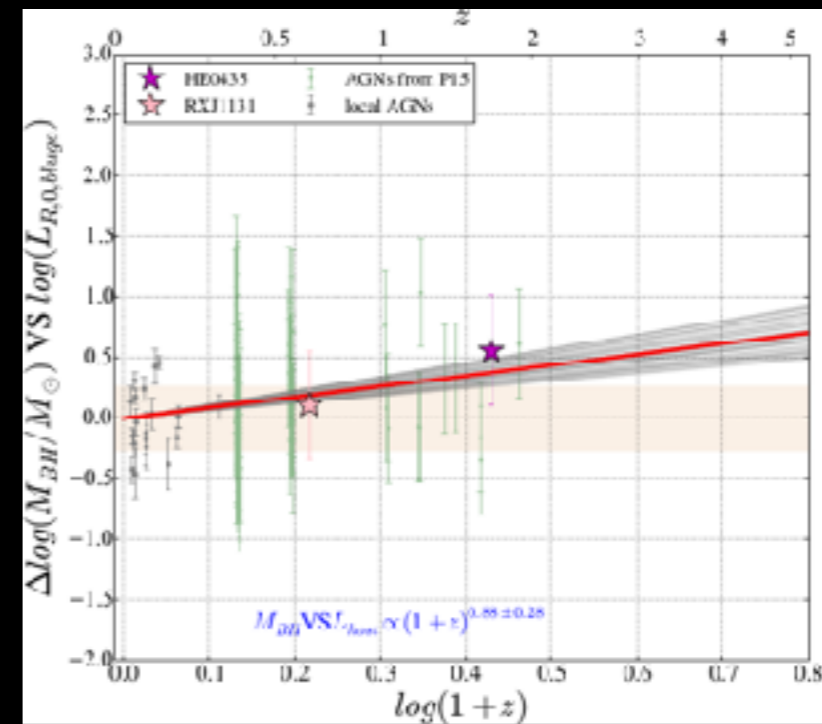
APRIM
7/4/2017

Ongoing Work

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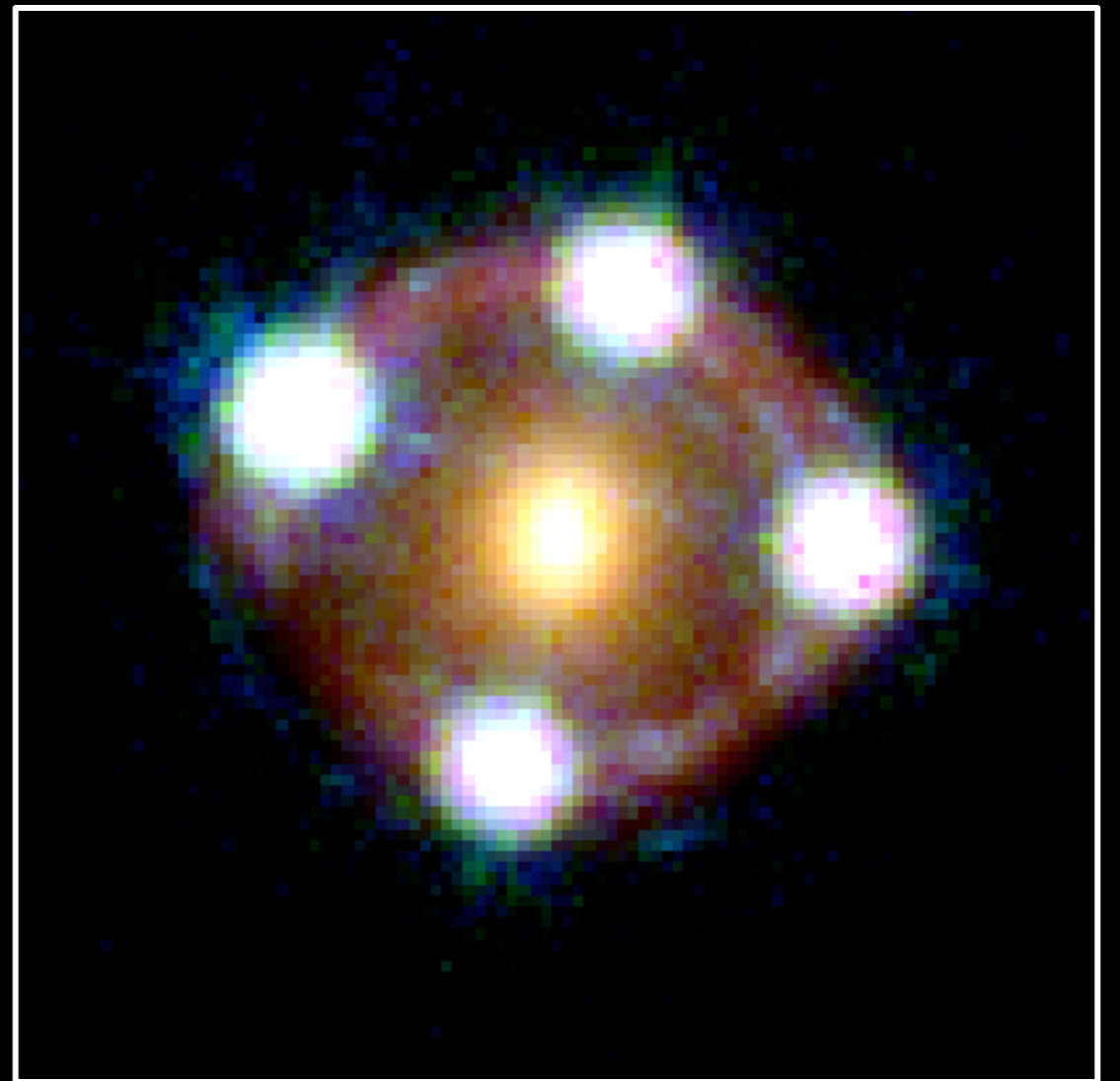
Ding+2017a



Ding+2017b

Summary

- Time-delay lenses are an independent probe of H_0
- Blind analysis of HE0435
 - time delays from COSMOGRAIL
 - deep *HST* imaging
 - wide-field imaging & spectroscopy
 - velocity dispersion from Keck/LRIS
- With 3 time-delay lenses from H0LiCOW: $H_0 = 71.9^{+2.4}_{-3.0}$ km/s/Mpc in flat Λ CDM
- Full H0LiCOW sample: H_0 to $< 3.5\%$ precision from 5 lenses (possibly $\sim 2\%$ precision from 9 lenses)
- Current and future surveys will find thousands of new time-delay lenses, providing competitive probe of cosmology



East Asian Young Astronomer's Meeting (EAYAM) 2017

Postdocs and grad students
from East Asia region
encouraged to participate

November 13-17, 2017
Ishigaki, Japan

Registration deadline: **July 15**

Go to the NAOJ booth here at
APRIM to pick up a flyer!



<http://www-irc.mtk.nao.ac.jp/eayam/2017/>



East Asia Core Observatories Association