

High spatial resolution polarimetric observations of hydroxyl maser emission towards the massive star-forming region G351.417+0.645

Thanapol Chanapote^{1,2,7}; Kitiyanee Asanok^{1,7}; Richard Dodson²; Maria Rioja^{2,3,4}; James Green⁵; Busaba Hutawarakorn Kramer⁶

¹*Department of Physics, Faculty of Science, Khon Kaen University (KKU), Khon Kaen, 40002, Thailand;* ²*International Centre for Radio Astronomy Research (ICRAR), The University of Western Australia (UWA), M468, 35 Stirling Highway, Crawley, Perth, WA 6009, Australia;* ³*CSIRO Astronomy and Space Science, 26 Dick Perry Avenue, Kensington WA 6151, Australia;* ⁴*Observatorio Astronómico Nacional (IGN), Alfonso XII, 3 y 5, 28014 Madrid, Spain;* ⁵*CSIRO Astronomy and Space Science, Australia Telescope National Facility, PO Box 76, Epping, NSW 1710, Australia;* ⁶*Max-Planck-Institut für Radioastronomie, Auf dem Hügel 69, 53121 Bonn, Germany;* ⁷*National Astronomical Research Institute Of Thailand (NARIT), Ministry of Science and Technology, Rama VI Rd., Bangkok 104000, Thailand*

In this study, we present high spatial resolution polarimetric observations obtained with the Long Baseline Array (LBA) of both ground and excited state hydroxyl (OH) maser emission towards the massive star-forming region G351.417+0.645. Our results provide the first well-characterized spatial gradient magnetic field measurements for the ground state transitions. These show the magnetic field strengths vary between -5.7 and +4.2 mG (negative and positive signs represent the magnetic field directed towards and away from us respectively), and comparably -5.7 to +4.4 mG for the excited state, with a clear spatial separation in the positive and negative field directions. This clear spatial separation, which coincides with a separation along the Galactic plane, will be explained in relation to the star forming region. Furthermore, the kinematics and environmental conditions in an 11 year time span will be examined through a comparison with previous LBA observations.